# Welcome back to 8.033!



1803-1853, Austrian

Image Courtesy of Wikipedia.

Why opposite sense?

**Summary of last lecture:** 

- Time dilation
- Length contraction
  - Relativity of simultaneity
  - Problem solving tips

# MIT Course 8.033, Fall 2006, Lecture 6 Max Tegmark

## **Today: Relativistic Kinematics**

- Space/time unification:  $\eta$ , imaginary rotations, etc.
- Proper time, rest length, timelike, spacelike, null
- More 4-vectors: U, K
- Velocity addition
- Doppler effect
- Aberration

# Velocity addition

#### Transformation toolbox: velocity addition

• If the frame S' has velocity  $v_1$  relative to S and the frame S" has velocity  $v_2$  relative to S' (both in the x-direction), then what is the speed  $v_3$  of S" relative to S?

• 
$$\mathbf{x}' = \mathbf{\Lambda}(v_1)\mathbf{x}$$
 and  $\mathbf{x}'' = \mathbf{\Lambda}(v_2)\mathbf{x}' = \mathbf{\Lambda}(v_2)\mathbf{\Lambda}(v_1)\mathbf{x}$ , so

•  $\Lambda(\mathbf{v}_3) = \Lambda(v_2)\Lambda(v_1), i.e.$ 

$$\begin{pmatrix} \gamma_3 & 0 & 0 & -\gamma_3\beta_3 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\gamma_3\beta_3 & 0 & 0 & \gamma_3 \end{pmatrix} = \begin{pmatrix} \gamma_2 & 0 & 0 & -\gamma_2\beta_2 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\gamma_2\beta_2 & 0 & 0 & \gamma_2 \end{pmatrix} \begin{pmatrix} \gamma_1 & 0 & 0 & -\gamma_1\beta_1 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\gamma_1\beta_1 & 0 & 0 & \gamma_1 \end{pmatrix}$$

$$= \gamma_1\gamma_2 \begin{pmatrix} 1+\beta_1\beta_2 & 0 & 0 & -[\beta_1+\beta_2] \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -[\beta_1+\beta_2] & 0 & 0 & 1+\beta_1\beta_2 \end{pmatrix}$$

• Take ratio between (1, 4) and (1, 1) elements:

$$eta_3=-rac{oldsymbol{\Lambda}(v_3)_{41}}{oldsymbol{\Lambda}(v_3)_{11}}=rac{eta_1+eta_2}{1+eta_1eta_2}.$$

• In other words,

$$v_3 = rac{v_1 + v_2}{1 + rac{v_1 v_2}{c^2}}.$$

SIMPLER WITH 2x2 MATRICES

#### Transformation toolbox: perpendicular velocity addition

- Here's an alternative derivation of velocity addition that easily gives the non-parallel components too
- If the frame S' has velocity v in the x-direction relative to S and a particle has velocity  $\mathbf{u}' = (u'_x, u'_y, u'_z)$  in S', then what is its velocity  $\mathbf{u}$  in S?
- Applying the inverse Lorentz transformation

$$egin{array}{rcl} x & = & \gamma(x'+vt') \ y & = & y' \ z & = & z' \ t & = & \gamma(t'+vx'/c^2) \end{array}$$

to two nearby points on the particle's world line and subtracting gives

$$egin{array}{rcl} dx&=&\gamma(dx'+vdt')\ dy&=&dy'\ dz&=&dz'\ dt&=&\gamma(dt'+vdx'/c^2). \end{array}$$

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• Answer:

$$\begin{aligned} u_x &= \frac{dx}{dt} = \frac{\gamma(dx' + vdt')}{\gamma(dt' + \frac{vdx'}{c^2})} = \frac{\frac{dx'}{dt'} + v}{1 + \frac{v}{c^2}\frac{dx'}{dt'}} = \frac{u'_x + v}{1 + \frac{u'_x v}{c^2}} \\ u_y &= \frac{dy}{dt} = \frac{dy'}{\gamma(dt' + \frac{vdx'}{c^2})} = \frac{\gamma^{-1}\frac{dy'}{dt'}}{1 + \frac{v}{c^2}\frac{dx'}{dt'}} = \frac{u'_y\sqrt{1 - \frac{v^2}{c^2}}}{1 + \frac{u'_x v}{c^2}} \\ u_z &= \frac{dz}{dt} = \frac{dz'}{\gamma(dt' + \frac{vdx'}{c^2})} = \frac{\gamma^{-1}\frac{dz'}{dt'}}{1 + \frac{v}{c^2}\frac{dx'}{dt'}} = \frac{u'_z\sqrt{1 - \frac{v^2}{c^2}}}{1 + \frac{u'_x v}{c^2}} \end{aligned}$$

# Unification of space & time

## "Everything is relative" — or is it?

- All observers agree on rest length
- All observers agree on proper time
- All observers (as we'll see later) agree on rest mass

Analogy: Australians and Bostonians agree on length of a 3vector.

#### Transformation toolbox: boosts as generalized rotations

- A "boost" is a Lorentz transformation with no rotation
- A rotation around the z-axis by angle  $\theta$  is given by the transformation

1	$\cos heta$	$\sin  heta$	0	0	١
	$-\sin heta$	$\cos heta$	0	0	
	0	0	1	0	
	0	0	0	1 ,	

• We can think of a boost in the x-direction as a rotation by an imaginary angle in the (x, ct)-plane:

$$oldsymbol{\Lambda}(-v) = \left(egin{array}{cccc} \gamma & 0 & 0 & \gammaeta \ 0 & 1 & 0 & 0 \ 0 & 0 & 1 & 0 \ \gammaeta & 0 & 0 & \gamma \end{array}
ight) = \left(egin{array}{cccc} \cosh\eta & 0 & 0 & \sinh\eta \ 0 & 1 & 0 & 0 \ 0 & 0 & 1 & 0 \ \sinh\eta & 0 & 0 & \cosh\eta \end{array}
ight),$$

where  $\eta \equiv \tanh^{-1} \beta$  is called the *rapidity*.

- Proof: use hyperbolic trig identities on next page
- Implication: for multiple boosts in same direction, rapidities add and hence the order doesn't matter

### Hyperbolic trig reminders

 $\cosh x = \frac{e^x + e^{-x}}{2}$  $\sinh x = \frac{e^x - e^{-x}}{2}$  $\tanh x = \frac{e^x - e^{-x}}{e^x + e^{-x}}$  $\cosh^{-1} x = \ln(x + \sqrt{x^2 - 1})$  $\sinh^{-1} x = \ln(x + \sqrt{x^2 + 1})$  $\tanh^{-1} x = \frac{1}{2} \ln \left( \frac{1+x}{1-x} \right)$  $\cosh \tanh^{-1} x = \frac{1}{\sqrt{1 - r^2}}$  $\sinh \tanh^{-1} x = \frac{x}{\sqrt{1-r^2}}$  $\cosh^2 x - \sinh^2 x = 1$ 

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# The Lorentz invariant

• The Minkowski metric

The only difference between space & time is a minus sign!

 $oldsymbol{\eta} = \left( egin{array}{cccccc} 1 & 0 & 0 & 0 \ 0 & 1 & 0 & 0 \ 0 & 0 & 1 & 0 \ 0 & 0 & 0 & -1 \end{array} 
ight)$ 

is left invariant by all Lorentz matrices  $\Lambda :$ 

 $\Lambda^t\eta\Lambda=\eta$ 

(indeed, this equation is often used to define the set of Lorentz matrices — for comparison,  $\Lambda^{t}I\Lambda = I$  would define rotation matrices)

• Proof: Show that works for boost along x-axis. Show that works for rotation along y-axis or z-axis. General case is equivalent to applying such transformations in succession.

• All Lorentz transforms leave the quantity

$$\mathbf{x}^t \boldsymbol{\eta} \mathbf{x} = x^2 + y^2 + z^2 - (ct)^2$$

invariant

• Proof:

$$\mathbf{x}^{\prime t} \boldsymbol{\eta} \mathbf{x}^{\prime} = (\mathbf{\Lambda} \mathbf{x})^{t} \boldsymbol{\eta} (\mathbf{\Lambda} \mathbf{x}) = \mathbf{x}^{t} (\mathbf{\Lambda}^{t} \boldsymbol{\eta} \mathbf{\Lambda}) \mathbf{x} = \mathbf{x}^{t} \boldsymbol{\eta} \mathbf{x}$$
  
(Also easy to see directly from top equation)

- (More generally, the same calculation shows that  $\mathbf{x}^t \boldsymbol{\eta} \mathbf{y}$  is invariant)
- So just as the usual Euclidean squared length  $|\mathbf{r}|^2 = \mathbf{r} \cdot \mathbf{r} = \mathbf{r}^t \mathbf{r} = \mathbf{r}^t \mathbf{I} \mathbf{r}$  of a 3-vector is rotaionally invariant, the generalized "length"  $\mathbf{x}^t \boldsymbol{\eta} \mathbf{x}$  of a 4-vector is Lorentz-invariant.
- It can be positive or negative

#### 4-vectors are null, spacelike or timelike:

• For events  $\mathbf{x}_1$  and  $\mathbf{x}_2$ , their Lorentz-invariant separation is defined as

$$\Delta \sigma^2 \equiv \Delta \mathbf{x}^t \boldsymbol{\eta} \Delta \mathbf{x} = \Delta x^2 + \Delta y^2 + \Delta z^2 - (c \Delta t)^2$$

- A separation  $\Delta \sigma^2 = 0$  is called *null*
- A separation  $\Delta \sigma^2 > 0$  is called *spacelike*, and

$$\Delta\sigma\equiv\sqrt{\Delta\sigma^2}$$

is called the *proper distance* (the distance measured in a frame where the events are simultaneous)

• A separation  $\Delta \sigma^2 < 0$  is called *timelike*, and

$$\Delta au \equiv \sqrt{-\Delta \sigma^2}$$

is called the *proper time interval* (the time interval measured in a frame where the events are at the same place)



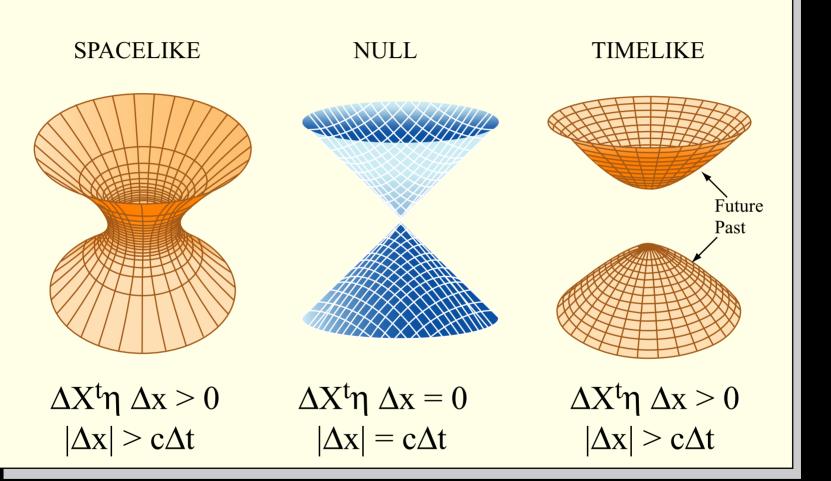


Figure by MIT OCW.

Timelike, spacelike or null?

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# Application:

# dt is invariant

#### Transformation toolbox: velocity as a 4-vector

• For a particle moving along its world-line, define its velocity 4-vector

$$\mathrm{U}\equiv rac{d\mathbf{X}}{d au}=\gamma_u \left(egin{array}{c} u_x\ u_y\ u_z\ c\end{array}
ight),$$

where

$$\gamma_u \equiv \frac{1}{\sqrt{1 - \frac{u^2}{c^2}}}$$

- This is the derivative of its 4-vector **x** w.r.t. its proper time  $\tau$ , since  $d\tau = dt/\gamma_u$
- $\mathbf{U}' = \mathbf{\Lambda}\mathbf{U}$ :

$$\mathbf{U}' = rac{d\mathbf{X}'}{d au'} = rac{d\mathbf{\Lambda}\mathbf{X}}{d au} = \mathbf{\Lambda}rac{d\mathbf{X}}{d au} = \mathbf{\Lambda}\mathbf{U},$$

since the proper time interval  $d\tau$  is Lorentz-invariant

• This means that all velocity 4-vectors are normalized so that

 $\mathbf{U}^t \boldsymbol{\eta} \mathbf{U} = -c^2$ . Which type?

• This immediately gives the velocity addition formulas:

where  $\gamma_{u'} = \gamma_u \gamma \left[1 + \frac{u_x v}{c^2}\right]$  — this last equation follows from the fact that the 4-vector normalization in Lorentz invariant, *i.e.*,  $\mathbf{u'}^t \eta \mathbf{u'} = \mathbf{u}^t \eta \mathbf{u} = -1$ .

• The 1st 3 components give the velocity addition equations we derived previously.

# This is how it *is* in the frame S'.

# But how does it *look*?

## Transforming a wave vector

• A plane wave

$$E(\mathbf{x}) = \sin(k_x x + k_y y + k_z z - \omega t)$$
(1)

is defined by the four numbers

$${f k}\equiv \left(egin{array}{c} k_x\ k_y\ k_z\ \omega/c\end{array}
ight)$$

- If the wave propagates with the speed of light c (like for an electromagnetic or gravitational wave), then the frequency is determined by the 3D wave vector  $(k_x, k_y, k_z)$  through the relation  $\omega/c = k$ , where  $k \equiv \sqrt{k_x^2 + k_y^2 + k_z^2}$
- How does the 4-vector **k** transform under Lorentz transformations? Let's see.

Using the Minkowski matrix, we can rewrite equation (1) as

 $E(\mathbf{x}) = \sin(\mathbf{k}^t \boldsymbol{\eta} \mathbf{x}).$ 

Let's Lorentz transform this:  $x\to x',\,k\to k'.$  Using that  $x'=\Lambda x,$  let's determine k'.

$$E' = \sin(\mathbf{k'}^t \boldsymbol{\eta} \mathbf{x'}) = \sin(\mathbf{k'}^t \boldsymbol{\eta} \mathbf{\Lambda} \mathbf{x}) = \sin[(\mathbf{\Lambda}^{-1} \mathbf{k'})^t (\mathbf{\Lambda}^t \boldsymbol{\eta} \mathbf{\Lambda}) \mathbf{x}] = \sin[(\mathbf{\Lambda}^{-1} \mathbf{k'})^t \boldsymbol{\eta} \mathbf{x}]$$

This equals E if  $\Lambda^{-1}\mathbf{k}' = \mathbf{k}$ , *i.e.*, if the wave 4-vector transforms just as a normal 4-vector:

$$\mathbf{k}' = \mathbf{\Lambda}\mathbf{k}$$

This argument assumed that E' = E. Later we'll see that the electric and magnetic fields do in fact change under Lorentz transforms, but not in a way that spoils the above derivation (in short, the phase of the wave,  $\mathbf{k}^t \boldsymbol{\eta} \mathbf{x}$ , must be Lorentz invariant)

So a plane wave k in S is also a plane wave in S', and the wave 4-vector transforms in exactly the same way as x does.

## Aberration and Doppler effects

• Consider a plane wave propagating with speed c in the frame S:

$$\mathbf{k}=k\left(egin{array}{c} \sin heta\cos\phi\ \sin heta\sin\phi\ \cos heta\ 1\end{array}
ight),$$

where ck is the wave frequency and the angles  $\theta$  and  $\phi$  give the propagation direction in polar coordinates.

Let's Lorentz transform this into a frame S' moving with speed v relative to S in the z-direction:  $\mathbf{k}' = \mathbf{\Lambda}\mathbf{k}$ , *i.e.*,

$$egin{array}{rcl} \mathbf{k}' &=& k' \left( egin{array}{c} \sin heta' \cos \phi' \ \sin heta' \sin \phi' \ \cos heta' \ 1 \end{array} 
ight) = k \left( egin{array}{ccc} 1 & 0 & 0 & 0 \ 0 & 1 & 0 & 0 \ 0 & 0 & \gamma & -\gamma eta \ 0 & 0 & -\gamma eta & \gamma \end{array} 
ight) \left( egin{array}{c} \sin heta \sin \phi \ \cos \phi \ 1 \end{array} 
ight) = k \left( egin{array}{c} \sin heta \cos \phi \ \sin heta \sin \phi \ \gamma(\cos heta - eta) \ \gamma(1 - eta \cos heta) \end{array} 
ight), \end{array}$$

 $\mathbf{SO}$ 

$$egin{array}{rcl} \phi'&=&\phi\ \cos heta'&=&rac{\cos heta-eta}{1-eta\cos heta}\ k'&=&k\gamma(1-eta\cos heta) \end{array}$$

- This matches equations (1)-(4) in the Weiskopf et al ray tracing handout
- The change in the angle  $\theta$  is known as *aberration*
- The change in frequency ck is known as the Doppler shift note that since  $k = 2\pi/\lambda$ , we have  $\lambda'/\lambda = k/k'$ .
- If we instead take the ratio  $\sqrt{k'_x^2 + k'_y^2}/k'_z$  above, we obtain the mathematically equivalent form of the aberration formula given by Resnick (2-27b):

$$an heta' = rac{\sin heta}{\gamma(\cos heta-eta)}$$

- Examine classical limits
- Transverse Doppler effect:  $\cos \theta = 0$  gives  $\omega' = \omega \gamma$ , *i.e.*, simple time dilation