### **Radiative Equilibrium**

- Equilibrium state of atmosphere and surface in the absence of non-radiative enthalpy fluxes
- Radiative heating drives actual state toward state of radiative equilibrium

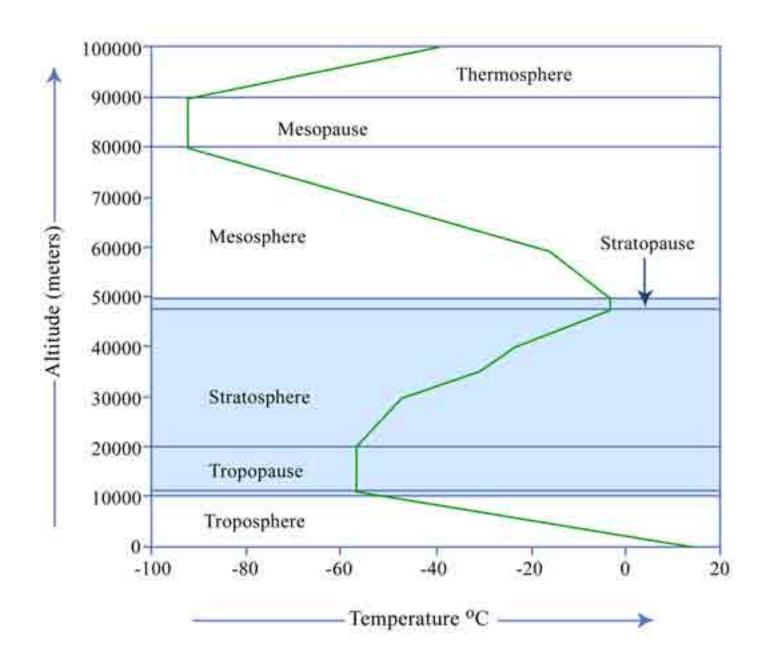
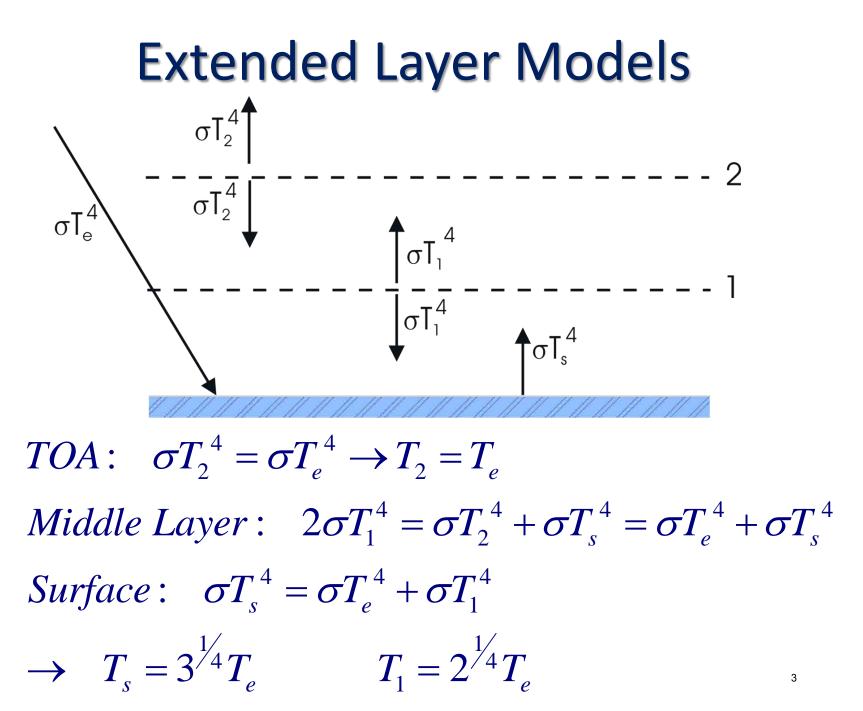
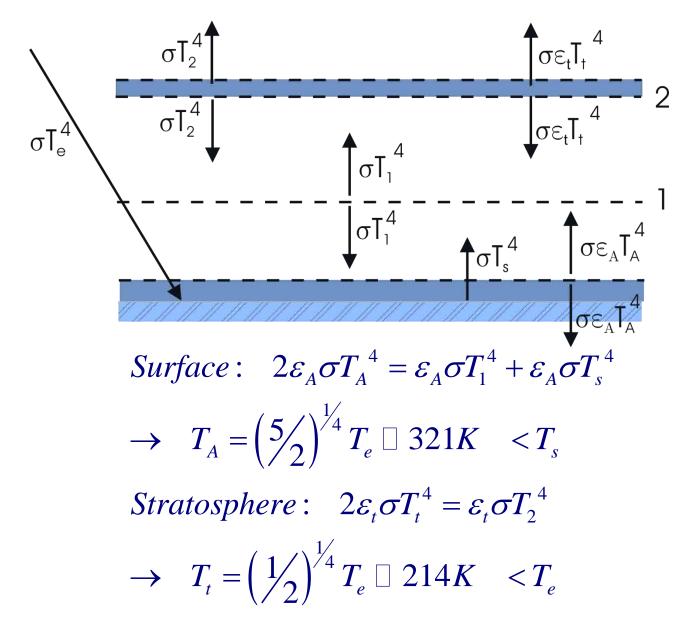


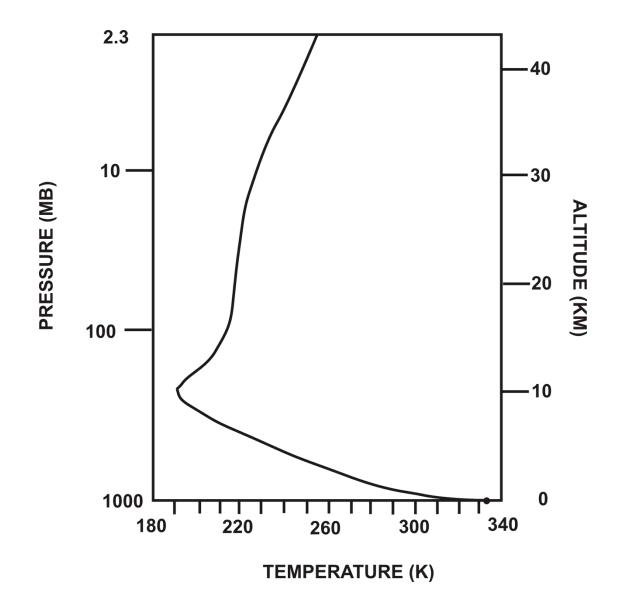
Figure by MIT OpenCourseWare.



### Effects of emissivity<1



#### Full calculation of radiative equilibrium



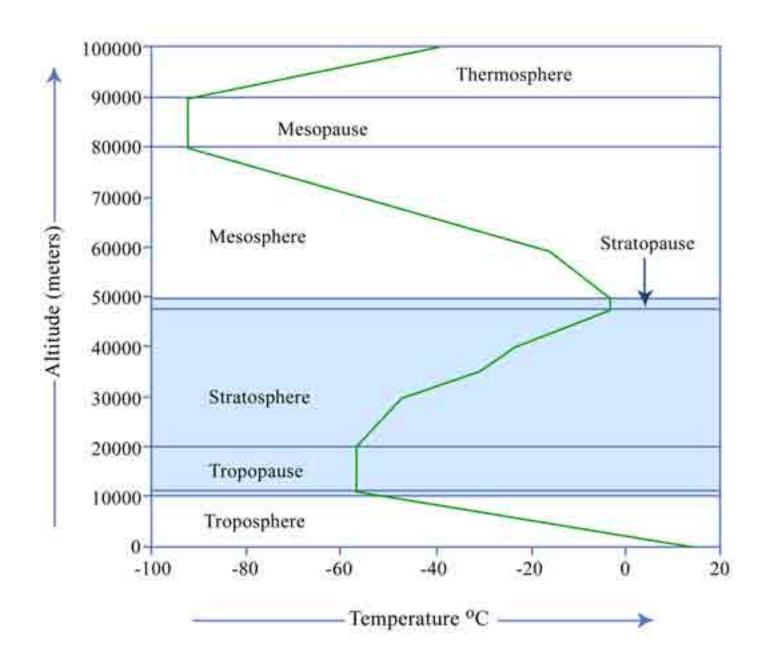
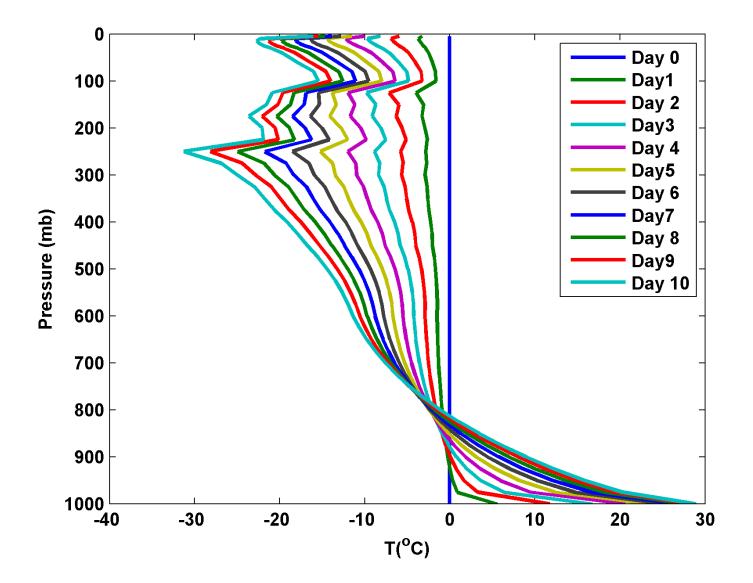
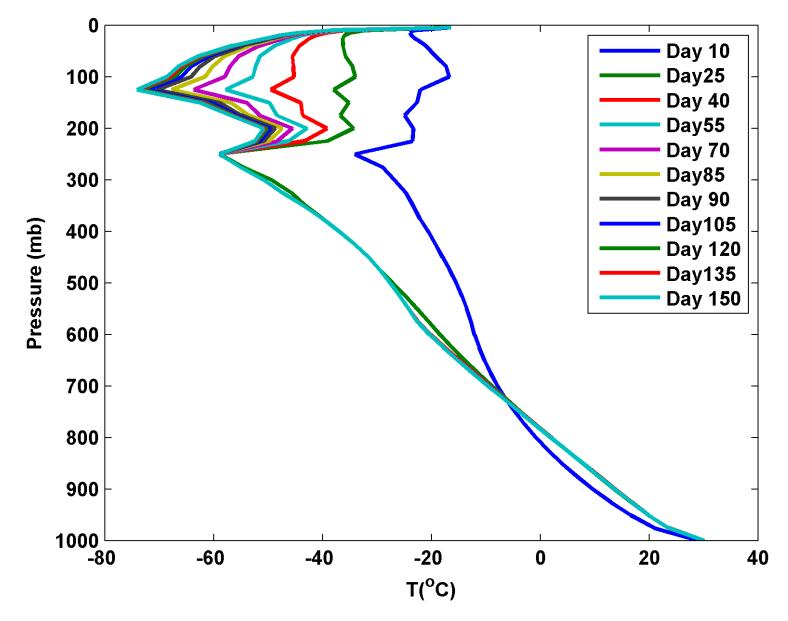


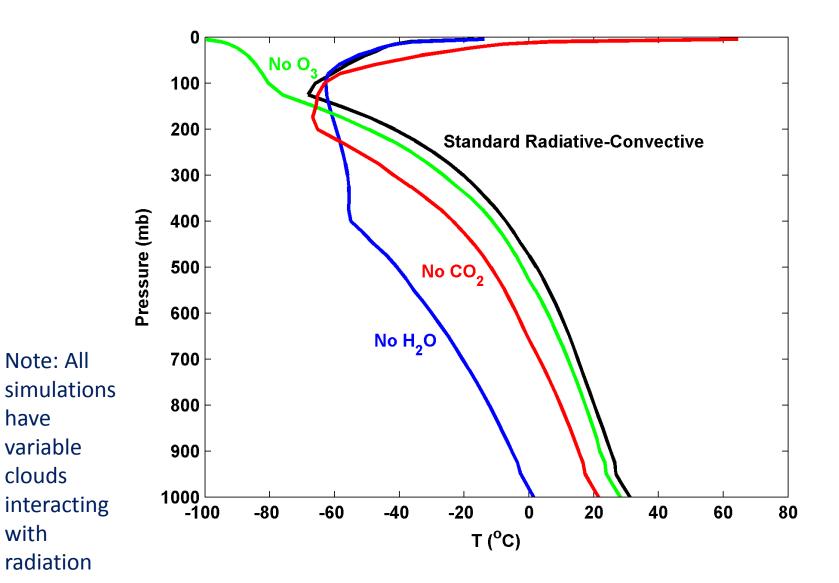
Figure by MIT OpenCourseWare.

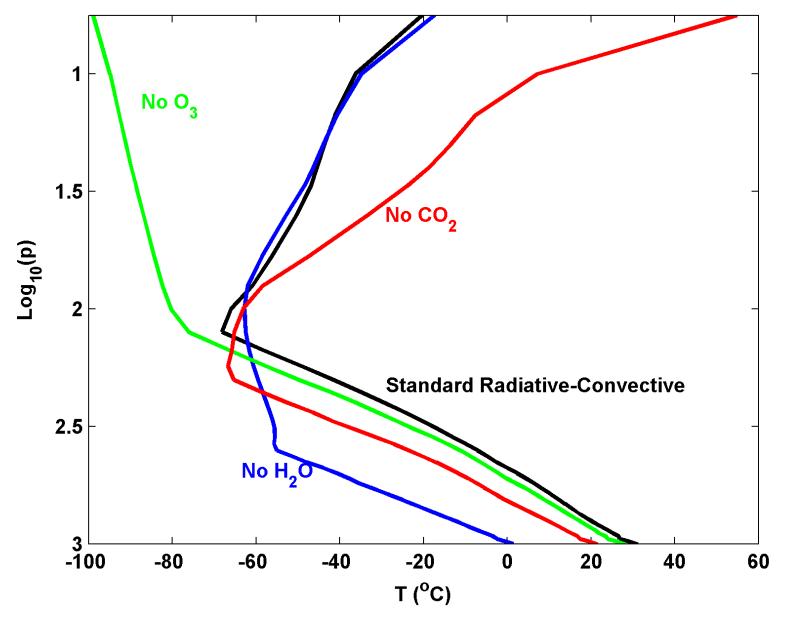
#### Time scale of approach to equilibrium





### **Contributions of various absorbers**



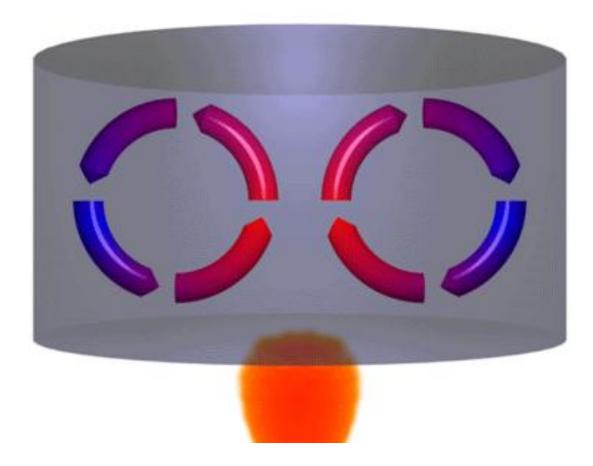


# Problems with radiative equilibrium solution

- Too hot at and near surface
- Too cold at a near tropopause
- Lapse rate of temperature too large in the troposphere
- (But stratosphere temperature close to observed)

### Missing ingredient: Convection

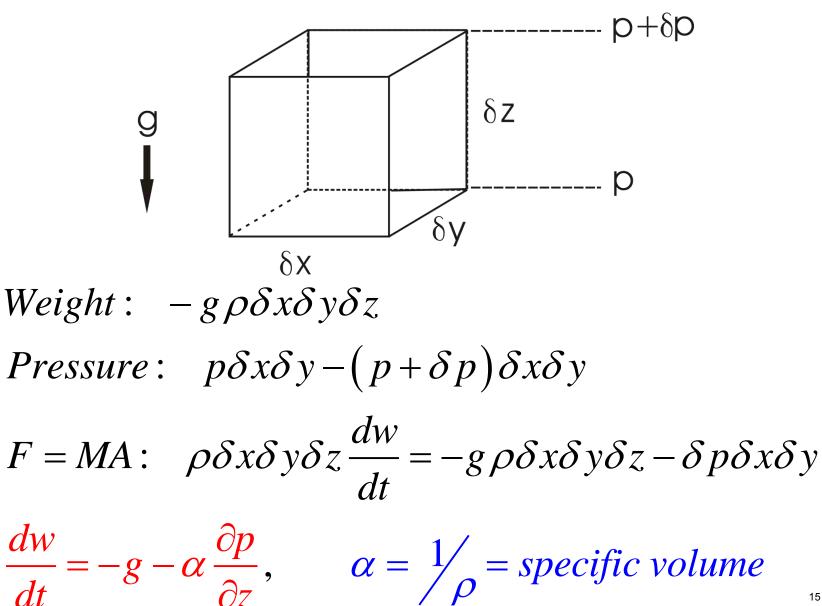
- As important as radiation in transporting enthalpy in the vertical
- Also controls distribution of water vapor and clouds, the two most important constituents in radiative transfer



# When is a fluid unstable to convection?

- Pressure and hydrostatic equilibrium
- Buoyancy
- Stability

### Hydrostatic equilibrium

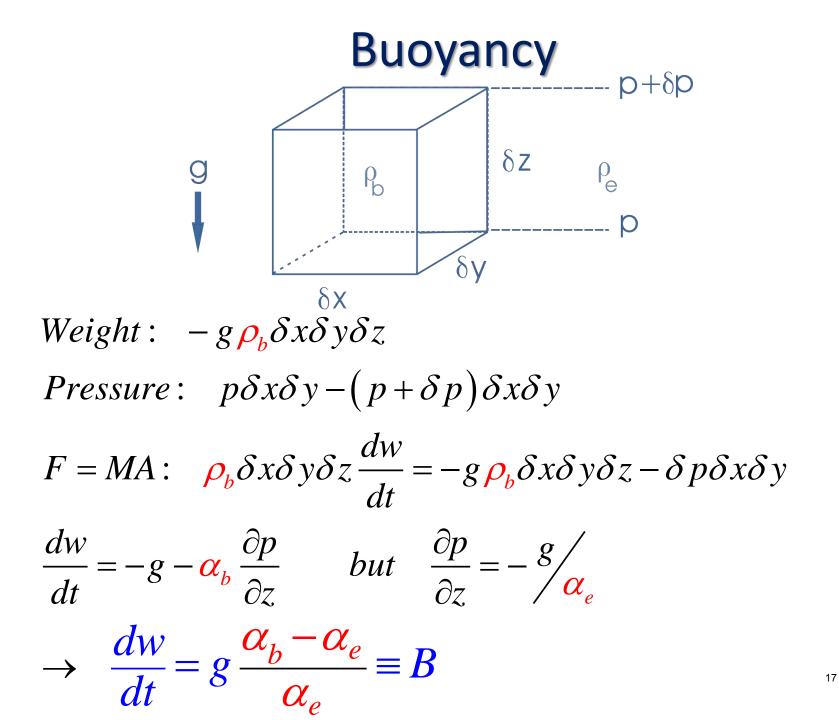


### Pressure distribution in atmosphere at rest

Ideal gas: 
$$\alpha = \frac{RT}{p}, \quad R \equiv \frac{R^*}{\overline{m}}$$
  
Hydrostatic:  $\frac{1}{p} \frac{\partial p}{\partial z} = -\frac{g}{RT}$ 

Isothermal case:  $p = p_0 e^{-\gamma_H}$ ,  $H \equiv \frac{\pi}{g} = "scale height"$ 

Earth: H~ 8 Km



### **Buoyancy and Entropy**

Specific Volume:

 $\alpha = \frac{1}{\rho}$ 

Specific Entropy:

 $\alpha = \alpha(p, s)$ 

S

 $\left(\delta\alpha\right)_p = \left(\frac{\partial\alpha}{\partial s}\right)_p \delta s = \left(\frac{\partial T}{\partial p}\right) \delta s$  $B = g \frac{\left(\delta \alpha\right)_p}{\alpha} = \frac{g}{\alpha} \left(\frac{\partial T}{\partial p}\right) \delta s = -\left(\frac{\partial T}{\partial z}\right)_s \delta s \equiv \Gamma \delta s$ 18

### The adiabatic lapse rate

First Law of Thermodynamics :

$$\dot{Q} = T \frac{ds_{rev}}{dt} = c_v \frac{dT}{dt} + p \frac{d\alpha}{dt}$$

$$= c_v \frac{dT}{dt} + \frac{d(\alpha p)}{dt} - \alpha \frac{dp}{dt}$$

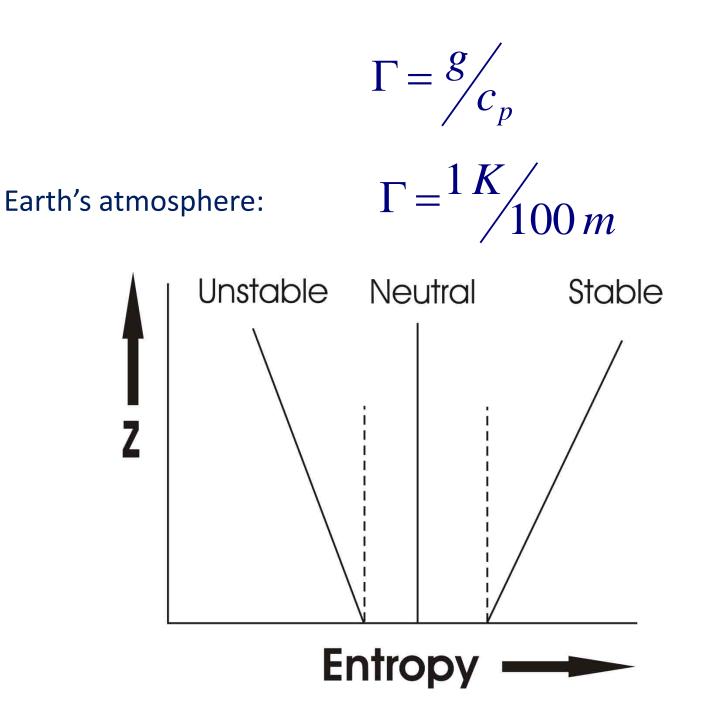
$$= (c_v + R) \frac{dT}{dt} - \alpha \frac{dp}{dt}$$

$$= c_p \frac{dT}{dt} - \alpha \frac{dp}{dt}$$

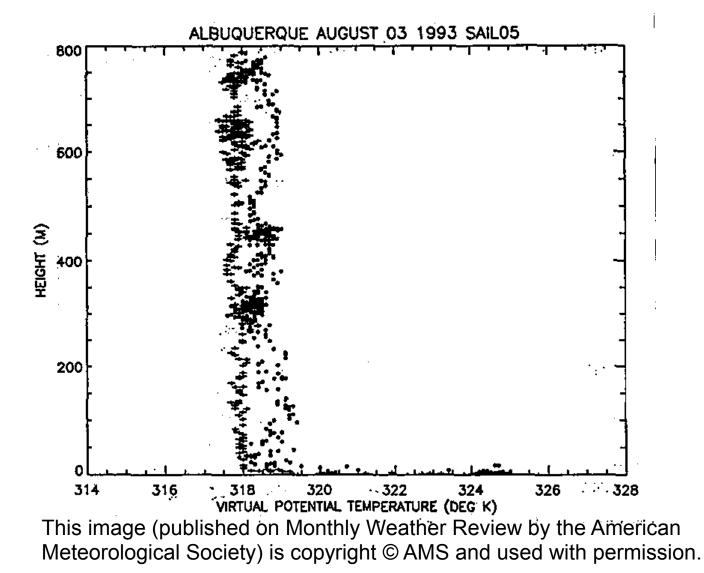
$$A diabatic: c_p dT - \alpha dp = 0$$

$$Hydrostatic: c_p dT + gdz = 0$$

$$\rightarrow (\frac{dT}{dz})_s = -\frac{g}{c_p} \equiv -\Gamma_d$$



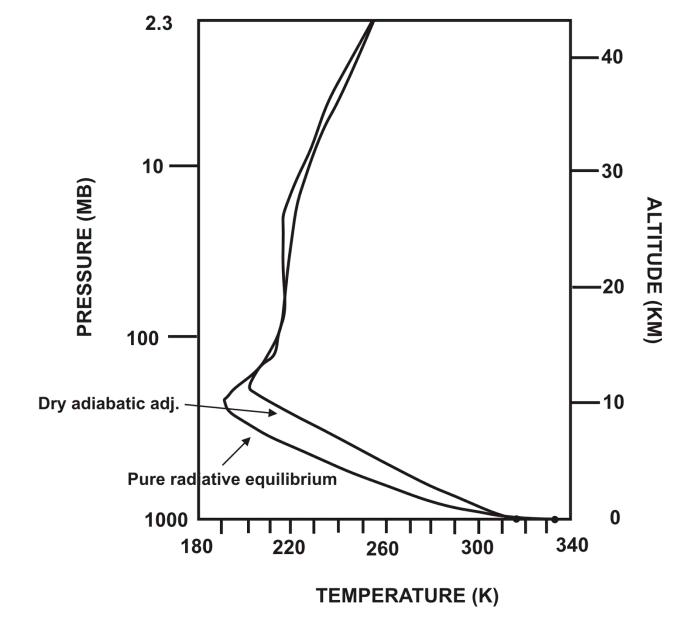
#### Model Aircraft Measurements (Renno and Williams, 1995)



### Radiative equilibrium is unstable in the troposphere

Re-calculate equilibrium assuming that tropospheric stability is rendered neutral by convection:

**Radiative-Convective Equilibrium** 



Better, but still too hot at surface, too cold at tropopause

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