

## LECTURE NOTE 7 \*

### RANDOM SAMPLE

MIT 14.30 SPRING 2006

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## 17 Definitions

### 17.1 Random Sample

Let  $X_1, \dots, X_n$  be mutually independent RVs such that  $f_{X_i}(x) = f_{X_j}(x) \forall i \neq j$ . Denote  $f_{X_i}(x) = f(x)$ . Then, the collection  $X_1, \dots, X_n$  is called a random sample of size  $n$  from the population  $f(x)$ .

Examples:

- Rolling a die  $n$  times.
- Selecting 10 MIT students and measuring their height.
- Sampling with and without replacement: Sampling from a large population (“nearly independent”).
- Alternatively, this collection (or sampling),  $X_1, \dots, X_n$ , is also called independent and identically distributed random variables with pmf/pdf  $f(x)$ , or iid sample for short.
- Note that the difference between  $X$  and  $x$  still holds (we continue to deal with random variables).

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\*Caution: These notes are not necessarily self-explanatory notes. They are to be used as a complement to (and not as a substitute for) the lectures.

## 17.2 Statistic

Let the RVs  $X_1, X_2, \dots, X_n$  be a random sample of size  $n$  from the population  $f(x)$ . Then, any real-valued function  $T = r(X_1, X_2, \dots, X_n)$  is called a statistic.

- Remember that  $X_1, X_2, \dots, X_n$  are RVs, and therefore  $T$  is a RV too, which can take any real value  $t$  with pmf/pdf  $f_T(t)$ .

## 17.3 Sample Mean

The sample mean, denoted by  $\bar{X}_n$ , is a statistic defined as the arithmetic average of the values in a random sample of size  $n$ .

$$\bar{X}_n = \frac{X_1 + X_2 + \dots + X_n}{n} = \frac{1}{n} \sum_{i=1}^n X_i \quad (52)$$

## 17.4 Sample Variance

The sample variance, denoted by  $S_n^2$ , is a statistic defined as:

$$S_n^2 = \frac{1}{n-1} \sum_{i=1}^n (X_i - \bar{X})^2 \quad (53)$$

The sample standard deviation is the statistic defined by  $S_n = \sqrt{S_n^2}$ .<sup>1</sup>

- Remember, the observed value of the statistic is denoted by lowercase letters. So,  $\bar{x}$ ,  $s^2$ , and  $s$  denote observed values of the RVs  $\bar{X}$ ,  $S^2$ , and  $S$ .

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<sup>1</sup>The sample variance and the sample standard deviation are sometimes denoted by  $\hat{\sigma}^2$  and  $\hat{\sigma}$ , respectively.

## 18 Important Properties of the Sample Mean Distribution and the Sample Variance Distribution

### 18.1 Mean and Variance of $\bar{X}$ and $S^2$

Let  $X_1, \dots, X_n$  be a random sample of size  $n$  from a population  $f(x)$  with mean  $\mu$  (finite) and variance  $\sigma^2$  (finite). Then,

$$E(\bar{X}) = \mu, \quad E(S^2) = \sigma^2, \quad \text{Var}(\bar{X}) = \frac{\sigma^2}{n}, \quad \text{and} \quad \text{Var}_{n \rightarrow \infty}(S^2) \rightarrow 0. \quad (54)$$

- Standard Error:  $\sqrt{\text{Var}(\bar{X})}$

**Example 18.1.** Show the first 3 statements of (54).

## 18.2 The Special Case of a Random Sample from a Normal Population

Let  $X_1, \dots, X_n$  be a random sample of size  $n$  from a  $N(\mu, \sigma^2)$  population. Then,

a.  $\bar{X}$  and  $S^2$  are independent random variables. (55)

b.  $\bar{X}$  has a  $N(\mu, \sigma^2/n)$  distribution. (56)

c.  $\frac{(n-1)S^2}{\sigma^2}$  has a  $\chi^2_{(n-1)}$  distribution. (57)

**Example 18.2.** Show (56).

## 18.3 Limiting Results ( $n \rightarrow \infty$ )

These concepts are extensively used in econometrics.

### 18.3.1 (Weak) Law of Large Numbers

Let  $X_1, \dots, X_n$  be independent and identically distributed (*iid*) random variables with  $E(X_i) = \mu$  (finite) and  $\text{Var}(X_i) = \sigma^2$  (finite). Define  $\bar{X}_n = \frac{1}{n} \sum_{i=1}^n X_i$ . Then, for every  $\varepsilon > 0$ ,

$$\lim_{n \rightarrow \infty} P(|\bar{X}_n - \mu| < \varepsilon) = 1. \quad (58)$$

This condition is denoted,

$$\bar{X}_n \xrightarrow{p} \mu \quad (\bar{X}_n \text{ converges in probability to } \mu.) \quad (59)$$

**Example 18.3.** Prove (58) using Chebyshev's inequality. Note that  $S^2 \xrightarrow{p} \sigma^2$  can be proved in a similar way.

### 18.3.2 Central Limit Theorem (CLT)

Let  $X_1, \dots, X_n$  be independent and identically distributed (*iid*) random variables with  $E(X_i) = \mu$  (finite) and  $\text{Var}(X_i) = \sigma^2$  (finite). Define  $\bar{X}_n = \frac{1}{n} \sum_{i=1}^n X_i$ . Then, for any value  $x \in (-\infty, \infty)$ ,

$$\lim_{n \rightarrow \infty} P\left(\frac{\sqrt{n}(\bar{X}_n - \mu)}{\sigma} < x\right) = \int_{-\infty}^x \frac{1}{\sqrt{2\pi}} e^{-x^2/2} = \Phi(x) \quad (60)$$

Where  $\Phi(\cdot)$  is the cdf of a standard normal.

*In words...* From (56) we know that if the  $X_i$ s are normally distributed, the sample mean statistic,  $\bar{X}_n$ , will also be normally distributed. (60) says that if  $n \rightarrow \infty$ , the function of the sample mean statistic,  $\frac{\sqrt{n}(\bar{X}_n - \mu)}{\sigma}$ , will be normally distributed **regardless** of the distribution of the  $X_i$ s.

*In practice(1)...* If  $n$  is sufficiently large, we can assume the distribution of a function of  $\bar{X}_n$ ,  $\frac{\sqrt{n}(\bar{X}_n - \mu)}{\sigma}$ , without knowing the underlining distribution of the random sample  $f_{X_i}(x)$ . [Very powerful result!]

In practice(2)...Define  $Z = \frac{\sqrt{n}(\bar{X}_n - \mu)}{\sigma}$ . If  $n$  is sufficiently large, then

$$F_Z \left( \frac{\sqrt{n}(\bar{x}_n - \mu)}{\sigma} \right) \approx \Phi \left( \frac{\sqrt{n}(\bar{x}_n - \mu)}{\sigma} \right) \quad (61)$$

↓

$$\frac{\sqrt{n}(\bar{X}_n - \mu)}{\sigma} \stackrel{a}{\sim} N(0, 1) \quad \text{or} \quad \bar{X}_n \stackrel{a}{\sim} N(\mu, \sigma^2/n) \quad (a : \text{for approximately}) \quad (62)$$

...regardless of the pmf/pdf  $f_{X_i}(x)$  !

- The larger the value of  $n$  is, the better the approximation. But, how much is “sufficiently large”? No straight forward rule. It will depend on the underlying distribution  $f_{X_i}(x)$ . The less bell-shaped  $f_{X_i}(x)$  is, the large the  $n$  required. Having said this, some authors suggest the following rule of thumb:  $n \geq 30$ .

- Magnifying glass (see simulations).

**Example 18.4.** An astronomer is interested in measuring the distance from his observatory to a distant star (in light years). Due to changing atmospheric conditions and measuring errors, each time a measurement is made it will not yield the exact distance. As a result, the astronomer plans to make several measures and then use the average as his estimated distance. He believes that measurement values are *iid* with mean  $d$  (the actual distance) and variance 4 (light years). How many measurements does he need to perform to be reasonably sure that his estimated distance is accurate to within  $\pm 0.5$  light years?