MIT OpenCourseWare http://ocw.mit.edu

16.346 Astrodynamics Fall 2008

For information about citing these materials or our Terms of Use, visit: http://ocw.mit.edu/terms.

## Lecture 12

Patched Conics and Planetary Alybys

 $r = \text{planet mean distance} \times \left(\frac{\text{planet mass}}{\text{solar mass}}\right)^{\frac{2}{5}}$ 

### Terminology

Sphere of Influence Radius r:

Inbound relative velocity vector:

Outbound relative velocity vector:

Turn angle  $2\nu$ :

Minimum passing distance:

Circular speed  $v_{\circ m}$  at radius  $r_m$ :

Point of aim  $r_a$ :

$\sin \nu =$	1	
	$1 + \frac{v_{\infty}^2}{v^2}$	
	$v_{\circ m}^2$	

# $\begin{aligned} \mathbf{v}_{\infty i} \\ \mathbf{v}_{\infty o} \\ \csc \nu &= e \\ v_{\infty}^{2} \sin 2\nu = |\mathbf{v}_{\infty o} \times \mathbf{v}_{\infty i}| \\ r_{m} &= q = a(1-e) = \frac{\mu}{v_{\infty}^{2}}(e-1) = \frac{\mu}{v_{\infty}^{2}}(\csc \nu - 1) \\ v_{om}^{2} &= \frac{\mu}{r_{m}} \\ r_{a} &= -ae \cos \nu = -a \cot \nu = r_{m} \sqrt{1 + 2\frac{v_{om}^{2}}{v_{\infty}^{2}}} \\ \hline \\ \overline{\tan \nu} &= \frac{\mu}{r_{a}v_{\infty}^{2}} \\ \hline \\ \hline \\ tion \text{ is } \end{aligned}$

### **Close Pass of a Target Planet**

The inbound plane of relative motion is the plane containing the orbital velocity vectors of the spacecraft and the planet. The unit normal to that plane is

 $\mathbf{i}_n = \text{Unit}(\mathbf{v}_{SC} \times \mathbf{v}_P)$ 

so that the vector point of aim is

$$\mathbf{r}_a = r_a \, \mathbf{i}_n \times \mathbf{i}_{\infty i}$$

The vector  $\mathbf{r}_a$  can be rotated about the direction of  $\mathbf{i}_{\infty i}$  so that the orientation of the passing plane, relative to the planet, can be controlled at will. The method for this is discussed in the following page.

 $r_m$  Planet  $r_a$  Inbound velocity of spacecraft v Orbital velocity of spacecraft v Orbital velocity of planet  $v_{\omega c}$ Inbound velocity  $v_{\omega c}$   $v_{\omega c}$ Fig. 9.1 from An Introduction to the Mathematics and Methods of Astrodynamics

Pages 419-423

Mathematics and Methods of Astrodynamics. Courtesy of AIAA. Used with permission.

Lecture 12

### The Rotation Matrix and the Rotation of a Vector

Let *O* be a fixed point on an axis whose direction is  $\mathbf{i}_{\omega} = l \, \mathbf{i}_x + m \, \mathbf{i}_y + n \, \mathbf{i}_z$  and let **r** be the vector from *O* to a point *P*. When the vector **r** rotates about this axis, the point *P* moves in a circle of radius  $r \sin \alpha$  where  $\alpha$  is the angle between **r** and  $\mathbf{i}_{\omega}$ .



Figure by MIT OpenCourseWare.

A small rotation through an angle  $\delta \psi$  produces a change  $\delta \mathbf{r}$  given by

$$\delta \mathbf{r} = \delta \psi \, \mathbf{i}_{\omega} \times \mathbf{r} = \delta \psi \, \mathbf{Sr}$$

where the constant matrix  ${\bf S}$  is skew-symmetric, i.e.,  ${\bf S}=-{\bf S}^{\, {\scriptscriptstyle \rm T}}$  . Hence

$$\frac{d\mathbf{r}}{d\psi} = \mathbf{Sr}$$

is a linear vector differential equation for  $\mathbf{r}$  with constant coefficients whose solution is

$$\mathbf{r} = e^{\psi \mathbf{S}} \mathbf{r}_0 \equiv \left( \mathbf{I} + \psi \mathbf{S} + \frac{1}{2!} \psi^2 \mathbf{S}^2 + \frac{1}{3!} \psi^3 \mathbf{S}^3 + \cdots \right) \mathbf{r}_0$$

Also, since

$$\mathbf{S} = \begin{bmatrix} 0 & -n & m \\ n & 0 & -l \\ -m & l & 0 \end{bmatrix} \quad \mathbf{S}^2 = \begin{bmatrix} -m^2 - n^2 & lm & ln \\ lm & -l^2 - n^2 & mn \\ ln & mn & -l^2 - m^2 \end{bmatrix} \quad \mathbf{S}^3 = \begin{bmatrix} 0 & n & -m \\ -n & 0 & l \\ m & -l & 0 \end{bmatrix}$$

we conclude that

$$S^3 = -S$$

Thus, all powers of the matrix  $\mathbf{S}$  are either  $\pm \mathbf{S}$  or  $\pm \mathbf{S}^2$  and we have

$$\mathbf{R} = \mathbf{I} + \left(\psi - \frac{1}{3!}\psi^3 + \cdots\right)\mathbf{S} + \left(\frac{1}{2!}\psi^2 - \frac{1}{4!}\psi^4 + \cdots\right)\mathbf{S}^2$$
$$\mathbf{R} = \mathbf{I} + \sin\psi\,\mathbf{S} + (1 - \cos\psi)\mathbf{S}^2$$

or

Hence  $\mathbf{r} = \mathbf{R}\mathbf{r}_0$  is equivalent to

$$\begin{aligned} \mathbf{r} &= \mathbf{r}_0 + \sin \psi \left( \mathbf{i}_{\omega} \times \mathbf{r}_0 \right) + \left( 1 - \cos \psi \right) \mathbf{i}_{\omega} \times \left( \mathbf{i}_{\omega} \times \mathbf{r}_0 \right) \end{aligned} \\ \text{Try this with } \mathbf{r}_0 &= \mathbf{i}_x \text{ and } \mathbf{i}_{\omega} = \mathbf{i}_z \end{aligned}$$

Application to Rotating the Point of Aim vector  $\mathbf{r}_a$  about the  $\mathbf{i}_{\infty i}$  direction

$$\mathbf{r}_{a} = r_{a} \mathbf{i}_{n} \times \mathbf{i}_{\infty i} \quad \text{or} \quad \mathbf{r}_{a} = \frac{\mu}{2v_{\infty}^{2} \sin^{2} \nu} \mathbf{i}_{\infty i} \times (\mathbf{i}_{\infty i} \times \mathbf{i}_{\infty o}) \qquad \text{See Problem 9-4}$$

$$\text{New } \mathbf{r}_{a} = \mathbf{r}_{a} + \sin \psi \left( \mathbf{i}_{\infty i} \times \mathbf{r}_{a} \right) + (1 - \cos \psi) \mathbf{i}_{\infty i} \times (\mathbf{i}_{\infty i} \times \mathbf{r}_{a})$$

# 16.346 Astrodynamics Lecture 12

### **Two Round-Trip Orbits**





# Round-Trip Earth to Venus to Earth Published August 1959

Round-Trip Earth to Venus to Mars to Earth Discovered January 1961





Planetary FlyBys are definitely three-dimensional

The normal to the plane in which the spacecraft moves in its hyperbolic swingby of the planet is determined by the cross-product of the inbound and outbound relative velocity vectors.



16.346 Astrodynamics

Lecture 12

The velocity components, which are part of this data, have never been published before. We give them below in units of feet per second:

Launch date: June 9, 19	972			
Outbound Earth velocity	$\begin{bmatrix} -6326 & -12628 & -5051 \end{bmatrix}$	Magnitude = $15000 \text{ Ft/sec}$		
Inbound Venus velocity	[28253  2698  335]	Magnitude = $28384$ Ft/sec		
Outbound Venus velocity	$[23564 \ 15771 \ -1274]$	Magnitude = $28383$ Ft/sec		
Inbound Mars velocity	[19624 - 17633 - 1790]	Magnitude = $26443$ Ft/sec		
Outbound Mars velocity	$\begin{bmatrix} -20418 & -16713 & 1736 \end{bmatrix}$	Magnitude = $26443$ Ft/sec		
Inbound Earth velocity	$\begin{bmatrix} -40370 & -7303 & 1836 \end{bmatrix}$	Magnitude = $41066 \text{ Ft/sec}$		
Return date: September 13, 1973				
Launch date: February	6, 1966			
Outbound Earth velocity	$\begin{bmatrix} -7907 & -13922 & 3985 \end{bmatrix}$	Magnitude = $16499$ Ft/sec		
Inbound Venus velocity	[29837  1164  -6313]	Magnitude = $30520$ Ft/sec		

Outbound Venus velocity	[23529  18335  -6453]	Magnitude = $30519$ Ft/sec
Inbound Mars velocity	[6061 - 14285 2957]	Magnitude = $15797$ Ft/sec
Outbound Mars velocity	$[8659 - 13164 \ 1115]$	Magnitude = $15796$ Ft/sec
Inbound Earth velocity	[39120  512  2839]	Magnitude = $39226$ Ft/sec
Return date: December	17, 1967	

Error correction in red was discovered by Steve Block (a student in 16.346) in the fall term of 2003 — over 42 years after the original posting!

The original velocity vector was incorrectly stated as  $\begin{bmatrix} 16713 & -20418 & 1736 \end{bmatrix}$ .

Lecture 12