

# Quiz 2

Name: (Last, First) \_\_\_\_\_ (please print).

Recitation number (circle one): 1 2 3

- Record all answers and show all work in this exam booklet. If you need extra space, use the back of the page.
- All scratch paper must be handed in with the exam, but will not be graded.
- This exam is closed book. You may use your handwritten notes if they are clearly labeled with your name and you hand them in with your exam.
- Whenever possible, try to solve problems using general analytic expressions. Plug in numbers only as a last step.
- Please make sure to answer all sub-questions.
- Good Luck!

Problem	Max	Grade	Grader
1	25		
2	25		
3	25		
4	25		
Total	100		

- (a). (5 pts) Consider the following reactions:
- (A) \_\_\_\_\_ A 100 MeV photon decays into an electron-positron pair. *answer:E*
  - (B) \_\_\_\_\_ A neutron decays into an electron-positron pair and a photon. *answer:B*
  - (C) \_\_\_\_\_ A neutron decays into a proton, an electron and a neutrino. *answer:L*
  - (D) \_\_\_\_\_ A proton decays into a neutron, a positron and a neutrino. *answer:E*
  - (E) \_\_\_\_\_ A neutron decays into a proton and a photon. *answer:Q*
- For each one, write **one** of the letters from the option list below.
- L violates lepton number conservation
  - B violates baryon number conservation
  - P violates parity conservation
  - E violates energy-momentum conservation
  - Q violates charge conservation
  - N violates none of the above conservation laws
- (b). (9 pts) Give each of the following quantities to the nearest power of 10 (don't show calculations, being off by one power of 10 is OK):
- (A) \_\_\_\_\_ Age of our universe when most He nuclei were formed *answer:1 min*
  - (B) \_\_\_\_\_ Age of our universe when hydrogen atoms formed. *answer:400000 yrs*
  - (C) \_\_\_\_\_ Age of our universe today. *answer:10 Gyr*
  - (D) \_\_\_\_\_ Number of stars in our Galaxy. *answer:1e11*
  - (E) \_\_\_\_\_ Light travel time to closest star (Sun!:) in minutes. *answer:8*
  - (F) \_\_\_\_\_ Hydrogen binding energy in  $\text{eV}/c^2$ . *answer:13.6*
  - (G) \_\_\_\_\_ Electron mass in  $\text{eV}/c^2$ . *answer:500000*
  - (H) \_\_\_\_\_ Neutron mass in  $\text{eV}/c^2$ . *answer:10<sup>9</sup>*
  - (I) \_\_\_\_\_ Light travel time to 2nd closest star in years. *answer:3*
- (c). (9 pts) Indicate whether each of the following statements are true or false.
- (A) TRUE / FALSE If our Universe is only X billion years old, then we can only see objects that are now less than X billion light years away *answer:F*
  - (B) TRUE / FALSE Space must be infinite, because it cannot end with a boundary without more space on the other side. *answer:F*
  - (C) TRUE / FALSE Leptons do not feel the strong interaction. *answer:T*
  - (D) TRUE / FALSE No experiment inside an isolated sealed lab in space can distinguish between whether it is uniformly accelerating or in a uniform gravitational field. *answer:T*
  - (E) TRUE / FALSE A clock by the ceiling runs faster than one by the floor. *answer:T*
  - (F) TRUE / FALSE Hubble's law implies that the Big Bang was an explosion localized near the comoving position of our Galaxy.
  - (G) TRUE / FALSE The expansion of our galaxy is governed by the Friedmann equation. *answer:F*
  - (H) TRUE / FALSE Two galaxies can recede from each other faster than the speed of light. *answer:T*
  - (I) TRUE / FALSE We know that our entire observable universe was once at infinite density *answer:F*
- (d). (2 pts) A tritium ( $\text{H}^3$ ) nucleus contains \_\_\_\_\_ up quarks and \_\_\_\_\_ down quarks. *answer:p+2n = 4 + 5*

In the Sun, one of the processes in the He fusion chain is  $p + p + e^- \rightarrow d + \nu$ , where  $d$  is a deuteron. Make the approximations that the deuteron rest mass is  $2m_p$ , and that  $m_e \approx 0$  and  $m_\nu \approx 0$ , since both the electron and the neutrino have negligible rest mass compared with the proton rest mass  $m_p$ .

- (a). For the arrangement shown in the figure, where (in the lab frame) the two protons have the same energy  $\gamma m_p$  and impact angle  $\theta$ , and the electron is at rest, calculate the energy  $E_\nu$  of the neutrino in the rest frame of the deuteron in terms of  $\theta$ ,  $m_p$  and  $\gamma$ .

*answer:* Use the fact that the quantity  $E^2 - p^2 c^2$  is invariant. In the deuteron's rest frame, after the collision:

$$E^2 - p^2 c^2 = (2m_p c^2 + E_\nu)^2 - E_\nu^2 \quad (2.1)$$

$$= 4m_p^2 c^4 + 4m_p c^2 E_\nu = 4m_p c^2 (m_p c^2 + E_\nu) \quad (2.2)$$

In the lab frame, before collision:

$$E^2 - p^2 c^2 = (2E_p)^2 - (2p_p \cos \theta c)^2 \quad (2.3)$$

$$= (2\gamma m_p c^2)^2 - (2\gamma \beta m_p \cos \theta c^2)^2 \quad (2.4)$$

Use  $\gamma^2 \beta^2 = (\gamma^2 - 1)$  in the second term and simplify the algebra to find

$$E^2 - p^2 c^2 = 4m_p^2 c^4 (\gamma^2 - (\gamma^2 - 1) \cos^2 \theta) \quad (2.5)$$

Equating the invariants in the two frames, we have

$$4m_p c^2 (m_p c^2 + E_\nu) = 4m_p^2 c^4 (\gamma^2 - (\gamma^2 - 1) \cos^2 \theta) \quad (2.6)$$

$$\Rightarrow E_\nu = m_p c^2 (\gamma^2 - 1) \sin^2 \theta \quad (2.7)$$

- (b). For the special case where the deuteron remains at rest in the lab frame and  $\theta = 30^\circ$ , solve for  $\gamma$  and calculate the energy of all particles (the deuteron, the neutrino, one of the protons) in terms of the proton rest mass  $m_p$ .

*answer:* The deuteron's rest frame is the lab frame. Also,  $\theta = 30^\circ$ . Use conservation of energy, along with the result from the previous part to find:

$$2\nu m_p c^2 = 2m_p c^2 + E_\nu \quad (2.8)$$

$$= 2m_p c^2 + m_p c^2 (\gamma^2 - 1)/4 \quad (2.9)$$

$$\Rightarrow 2\gamma = 2 + \gamma^2/4 - 1/4 \quad (2.10)$$

$$\Rightarrow \gamma = 7, 1 \quad (2.11)$$

$\gamma = 1$  is obviously not the solution. Thus,  $\gamma = 7$  and the energies are:  $E_p = 7m_p$ ,  $E_\nu = 12m_p$ ,  $E_d = 2m_p$

**Question 3: Coulomb's Law generalized**

[25 Points]

In an inertial frame  $S$ , the position  $\mathbf{r}_q$  of a point charge  $q$  moves according to  $\mathbf{r}_q(t) = v\hat{\mathbf{z}}t$ , i.e. with velocity  $v$  in the  $\hat{\mathbf{z}}$ -direction, passing the origin at  $t = 0$ . In the moving frame  $S'$  where the charge is at rest at the origin, Coulomb's law states that the electric field is

$$\mathbf{E}' = A \frac{\mathbf{r}'}{r'^3},$$

where  $A = q/4\pi\epsilon_0$ . Show that in the frame  $S$ , the electric field at  $t = 0$  is

$$\mathbf{E} = A \frac{(1 - \beta^2)}{(1 - \beta^2 \sin^2 \theta)^{3/2}} \frac{\mathbf{r}}{r^3},$$

where  $\theta$  is the usual polar angle ( $z = r \cos \theta$ ,  $x^2 + y^2 = r^2 \sin^2 \theta$ ).

*answer:* Let us convert all quantities to the cartesian coordinates. In the frame  $S'$ , the components of the electric field are:

$$\vec{E}' = E'_x \hat{x}' + E'_y \hat{y}' + E'_z \hat{z}' \quad (3.1)$$

$$E'_x = \frac{Ax'}{(x'^2 + y'^2 + z'^2)^{3/2}} \quad (3.2)$$

$$E'_y = \frac{Ay'}{(x'^2 + y'^2 + z'^2)^{3/2}} \quad (3.3)$$

$$E'_z = \frac{Az'}{(x'^2 + y'^2 + z'^2)^{3/2}} \quad (3.4)$$

We can now Lorentz transform the fields and coordinates from  $S'$  to  $S$ . First the coordinates,

$$x = x' \quad (3.5)$$

$$y = y' \quad (3.6)$$

$$\gamma z = z' \quad (3.7)$$

and then the fields,

$$E_x = \gamma E'_x = \frac{A\gamma x}{(x^2 + y^2 + \gamma^2 z^2)^{3/2}} \quad (3.8)$$

$$E_y = \gamma E'_y = \frac{A\gamma y}{(x^2 + y^2 + \gamma^2 z^2)^{3/2}} \quad (3.9)$$

$$E_z = E'_z = \frac{A\gamma z}{(x^2 + y^2 + \gamma^2 z^2)^{3/2}} \quad (3.10)$$

Note that the primed coordinates have been converted to the unprimed ones using the coordinate transformation. The total magnitude for the electric field in the  $S$  frame can be obtained from

$$E^2 = E_x^2 + E_y^2 + E_z^2 \quad (3.11)$$

$$= \frac{A^2 \gamma^2 r^2}{(x^2 + y^2 + \gamma^2 z^2)^3} \quad (3.12)$$

$$= \frac{A^2 r^2}{(1 - \beta^2)(x^2 + y^2 + z^2/(1 - \beta^2))^3} \quad (3.13)$$

$$\Rightarrow E = \frac{A(1 - \beta^2)}{r^2(1 - \beta^2 \sin^2 \theta)^{3/2}} \quad (3.14)$$

Since the electric field always has to be radial,  $\vec{E} = |E|\hat{r}$ .

- (a). **(10 pts)** Consider a particle coasting in the  $r$ -direction (i.e., with constant  $\theta$  and  $\phi$ ) in a flat FRW metric, with no non-gravitational forces acting on it. Use variational calculus to prove that  $p \propto 1/a$  (here  $p = m_0\gamma u$  is its momentum and  $u \equiv a\dot{r}$  is its velocity relative to nearby comoving observers). *answer:*The particle only has a radial motion  $\Rightarrow d\theta = d\phi = 0$ . Also, the universe is flat  $\Rightarrow k = 0$ . Thus, the FRW metric becomes:

$$d\tau^2 = dt^2 - a^2 dr^2 \tag{4.1}$$

$$\Rightarrow \Delta\tau = \int d\tau = \int \sqrt{dt^2 - a^2 dr^2} \tag{4.2}$$

$$= \int dt \sqrt{1 - a^2 \dot{r}^2} = \int dt f(t, r, \dot{r}) \tag{4.3}$$

The interval  $\Delta\tau$  has to be extremized. The Euler lagrange equations give:

$$\frac{\partial f}{\partial r} - \frac{d}{dt} \left[ \frac{\partial f}{\partial \dot{r}} \right] = 0 \tag{4.4}$$

$$\Rightarrow \frac{d}{dt} \left[ \frac{a^2 \dot{r}}{\sqrt{1 - a^2 \dot{r}^2}} \right] = 0 \tag{4.5}$$

$$\Rightarrow \frac{a^2 \dot{r}}{\sqrt{1 - a^2 \dot{r}^2}} = \text{constant} \tag{4.6}$$

Identifying  $a\dot{r} = u$  leave us with  $\gamma u \propto 1/a \Rightarrow p = m_0\gamma u \propto 1/a$ .

- (b). **(2 pts)** Given a), the value of  $u$  in the limit  $a \rightarrow 0$  is \_\_\_\_\_. *answer:*1(c)  
 (c). **(2 pts)** Given a), the value of  $u$  in the limit  $a \rightarrow \infty$  is \_\_\_\_\_. *answer:*0  
 (d). **(2 pts)** Thus relative to comoving observers, your results show that an object without external forces in an expanding universe **(circle one)** REMAINS IN UNIFORM MOTION / SLOWS DOWN / ACCELERATES. *answer:*slows down  
 (e). **(3 pts)** Starting with the answer from a), derive how the wavelength  $\lambda$  of a photon depends on  $a$ . Your answer should be of the form  $\lambda \propto$ (function of  $a$ ). *answer:*For a photon,  $p = hk/2\pi \propto 1/\lambda$ . So  $\lambda \propto 1/p \propto a$ .  
 (f). **(6 pts)** Solve the Friedmann equation

$$H^2 = \frac{8\pi G}{3} \rho - \frac{kc^2}{a^2}$$

to obtain a solution of the form  $a(t) \propto$  (function of  $t$ ) for the case where space is flat and the density is dominated by photons, and compute the age of the universe at the time when  $H^{-1} = 30$  seconds.

*answer:*For a flat universe dominated by photons,

$$\left( \frac{\dot{a}}{a} \right)^2 = H^2 \propto a^{-4} \tag{4.7}$$

$$\Rightarrow \dot{a} \propto a^{-1} \tag{4.8}$$

$$\Rightarrow a = At^{1/2} \tag{4.9}$$

$$\Rightarrow H = \dot{a}/a = 1/2t \tag{4.10}$$

Thus,  $t = H^{-1}/2 = 15\text{seconds}$ .