

Readings for today: Section 1.3 – Atomic Spectra, Section 1.7 up to equation 9b – Wavefunctions and Energy Levels, Section 1.8 – The Principle Quantum Number. (Same sections in 5th and 4th ed.)

Read for Lecture #6: Sections 1.9, 1.10, and 1.11 (Same sections in 4th ed.)

Assignment: Problem set #2 (due Thursday, September 18th at 5 pm).

Topics: The Schrödinger Equation and Hydrogen Atom Energy Levels

I. Binding energies of the electron to the nucleus for a hydrogen atom

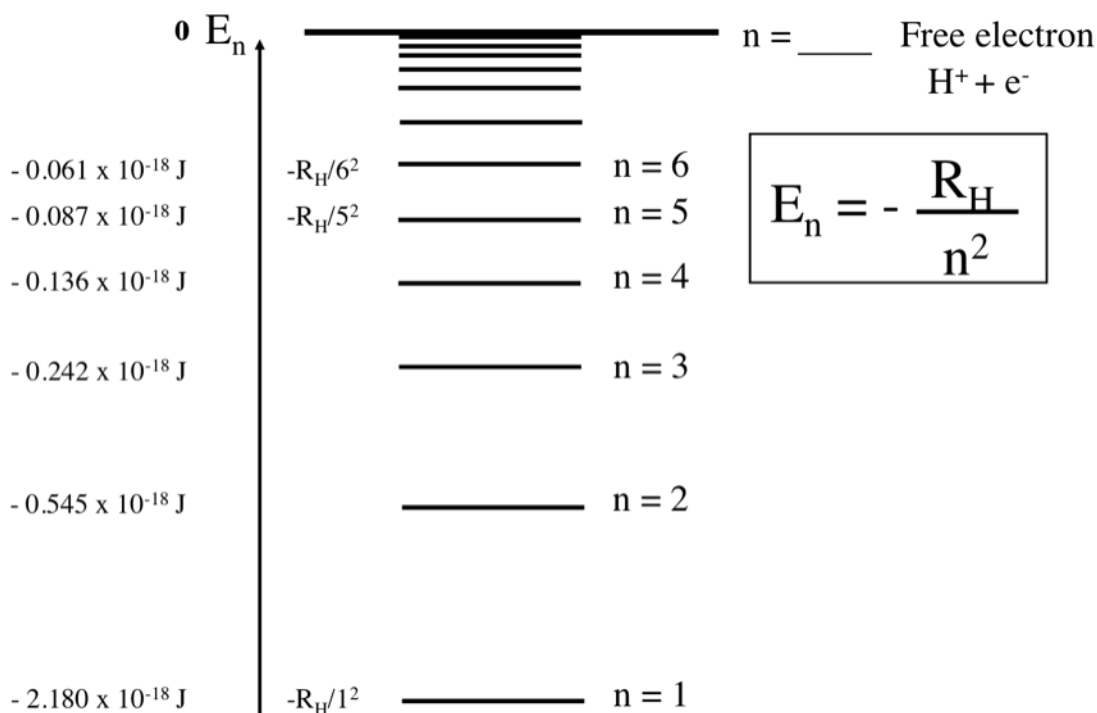
II. Verification of hydrogen-atom energy levels

A. Photon emission

B. Photon absorption

The Schrödinger equation is an equation of motion for particles (like electrons) that accounts for their wave-like properties. Solutions to the Schrödinger equation indicate possible binding energies and wavefunctions.

Energy level diagram for the H atom



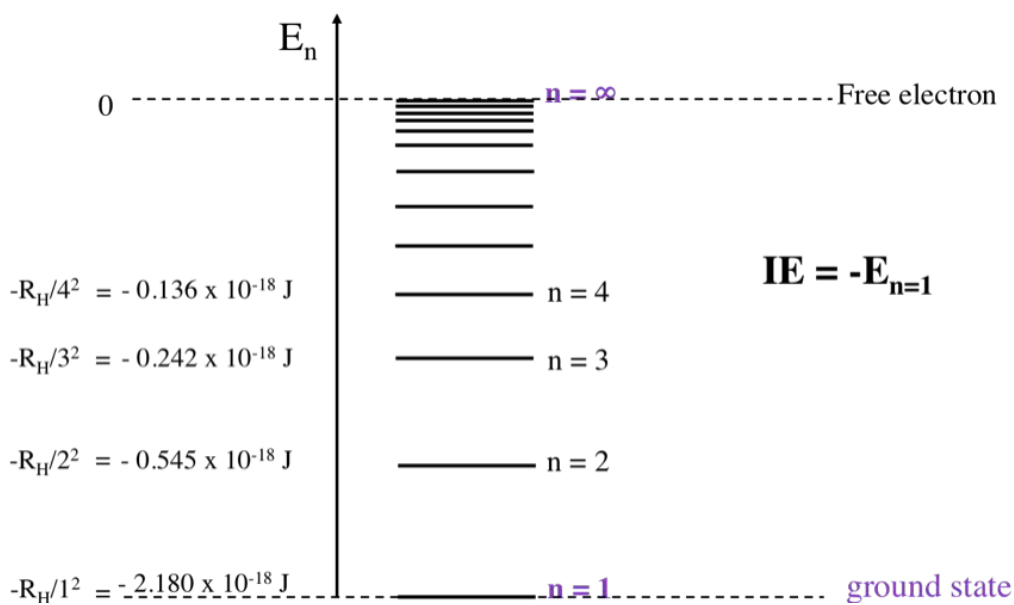
Note that all binding energies are _____ . A negative value means that the electron is bound to the nucleus. At $n=\infty$ ($E_n = 0$), the e^- is free from the nucleus.

The lowest (most negative) energy is called the _____ .

- The ground state is the most stable state.
- The ground state is the $n = 1$ state.

Ionization energy (IE) is the minimum energy required to remove an electron from the n^{th} state of a gaseous atom, molecule or ion. (Assume ground state, $n=1$, unless otherwise specified.)

- $E_n =$ _____ (ionization energy) of the hydrogen atom in the n^{th} state.
- Ionization energy is always _____. You always need to put energy into a system to eject an electron.
- The IE for a hydrogen atom in the ground state = _____ J. This means if you put that amount of energy into a hydrogen atom in its ground state, the electron is no longer bound to the nucleus.
- The IE for a hydrogen atom in the $n = 2$ (first excited state) is _____ J.
- The IE of a hydrogen atom in the **third** excited state ($n = __$) is _____ J.



The following equation describes the binding energy for any one-electron atom (including ions):

$$E_n = - \frac{R_H Z^2}{n^2} \quad \text{where } Z = \text{atomic number}$$

Electron is more weakly bound when n is big and more tightly when Z is big.

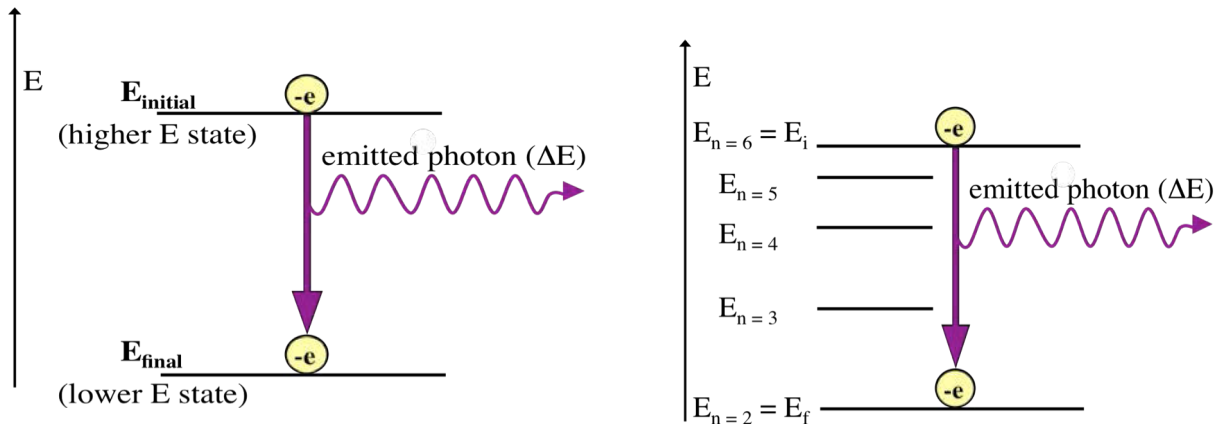
Atoms or ions with one electron:

- | | | |
|-------------------|---------------------|-------------------------|
| H | ≡ one electron atom | $Z = 1$ (atomic number) |
| He ⁺ | ≡ one electron ion | $Z = 2$ |
| Li ²⁺ | ≡ one electron ion | $Z = ______$ |
| Tb ⁶⁴⁺ | ≡ one electron ion | $Z = ______$ |

II. VERIFICATION OF HYDROGEN ATOM ENERGY LEVELS

A. PHOTON EMISSION

Photon emission occurs when an excited H atom relaxes to a lower E state. As the electron transitions from the higher to the lower E state, a photon is emitted that has the _____ energy as the energy difference between the two states.



We can calculate the energy of the emitted photon

$$\Delta E = \text{---} - \text{---}$$

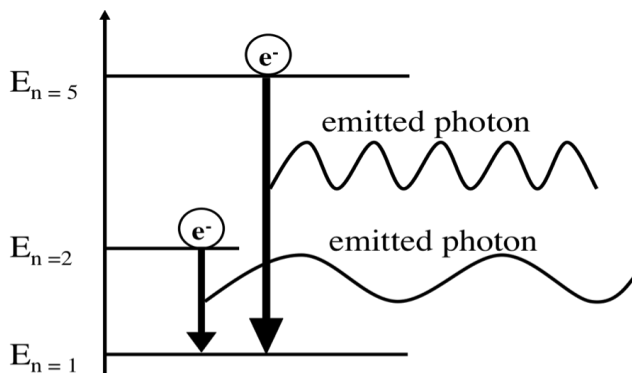
$$\Delta E = \text{---} - \text{---}$$

From the energy of the emitted photon, we can also calculate the frequency (ν) using $\Delta E = h\nu$.

$$\nu = \text{---}$$

$$\nu = \text{---}$$

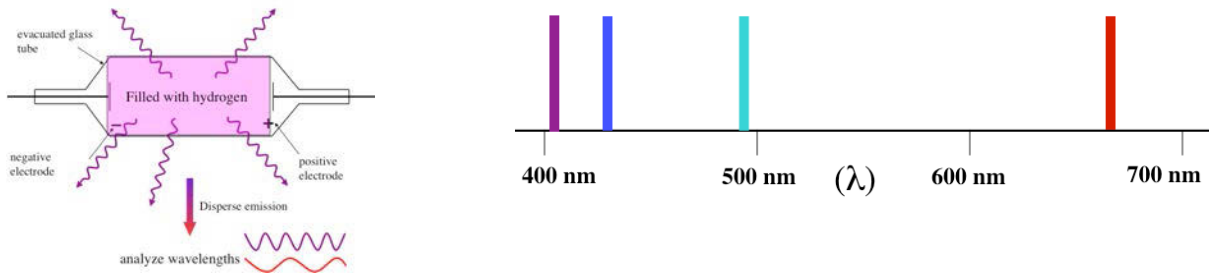
Consider the relationship between E, λ , and ν for any emitted photon:



$$n_i=5 \Rightarrow n_f=1 \text{ large } \Delta E \text{ --- } \nu \text{ --- } \lambda$$

$$n_i=2 \Rightarrow n_f=1 \text{ small } \Delta E \text{ --- } \nu \text{ --- } \lambda$$

Demonstration: Observing spectral lines from the visible spectrum of atomic hydrogen.



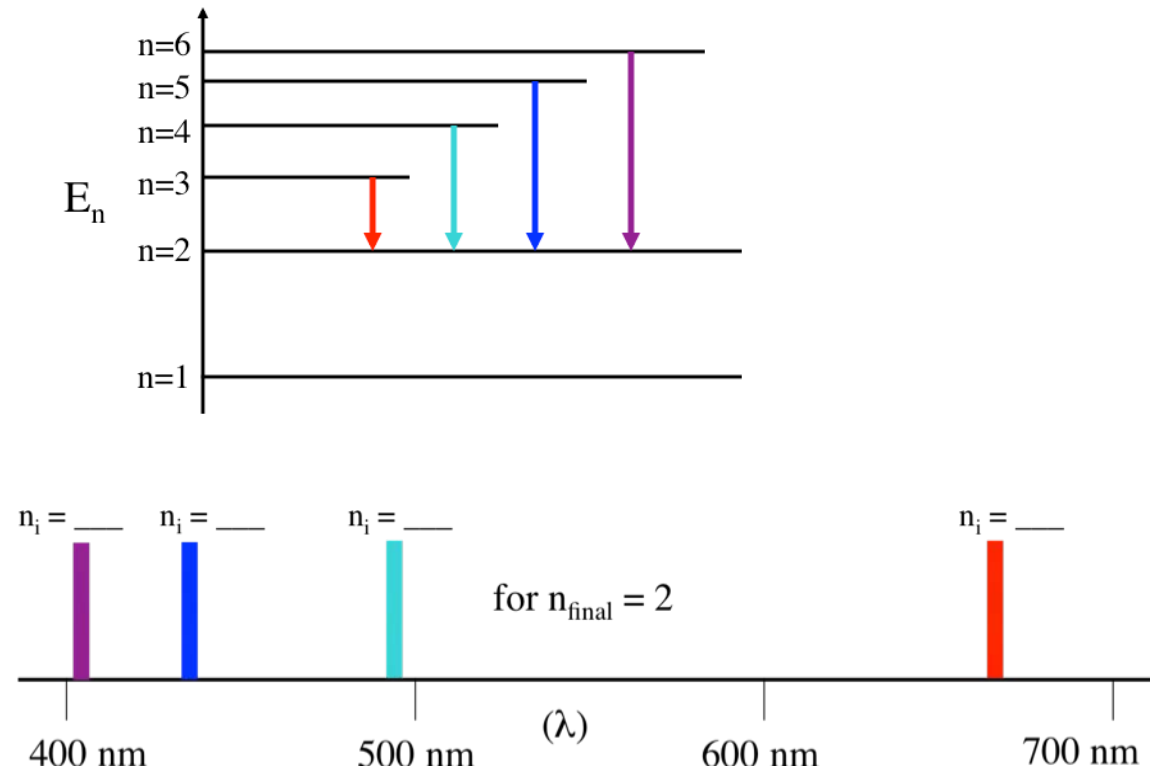
We are far from the first to observe these lines from atomic hydrogen!

1885 J.J. Balmer observed H atoms emit a series of lines in the visible region whose frequencies can be described by a simple formula:

$$\nu = 3.29 \times 10^{15} \text{ s}^{-1} \left[\frac{1}{4} - \frac{1}{n^2} \right] \quad \text{where } n = 3, 4, 5 \text{ or } 6$$

The origin of this formula was not understood at the time, but we now know:

- The lines result from electron transitions with a final energy level of $n = 2$.
- The frequency values can be accurately calculated using the relationship $E = h\nu$.



For the visible lines in the spectrum of atomic hydrogen, $E_f = E_{n=2}$.

We can calculate the predicted frequency and wavelength of these transitions.

$$\nu = \frac{E_i - E_f}{h}$$

and from the solution to the Schrödinger equation, we know

$$E_n = \frac{-R_H}{n^2}$$

So, $\nu = \frac{R_H}{h} \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$

$$\nu = \frac{R_H}{h} \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

For $n_f = 2$, then

$$\nu = \frac{R_H}{h} \left(\frac{1}{2^2} - \frac{1}{n_i^2} \right) \quad \text{BALMER SERIES}$$

$$R_H/h = \mathfrak{R} = 3.2898 \times 10^{15} \text{ s}^{-1}.$$

Once ν is calculated for $n_i = 3, 4, 5, 6 \dots$ use $\lambda = c/\nu$ to calculate λ .

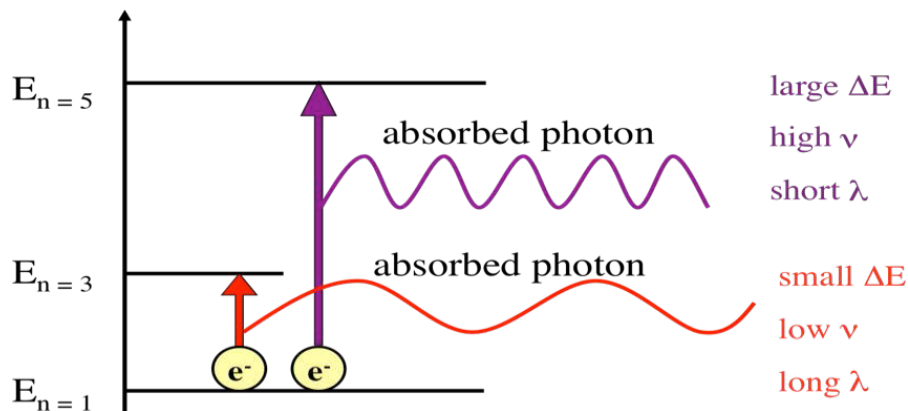
Calculations for λ using this equation derived from values of E_n predicted by Schrödinger equation match the observed λ or ν of emission to one part in 10^8 !

Transitions made to all final states from higher lying states. Named series include:

$n_f = 1$	Lyman series	_____
$n_f = 2$	Balmer series	visible range
$n_f = 3$	Paschen series	_____
$n_f = 4$	Brackett series	_____

B. PHOTON ABSORPTION

Energy can also be absorbed, exciting an electron to a higher energy state.



The frequency of the light absorbed can be calculated using:

$$\boxed{\nu = \frac{R_H}{h} \left(\frac{1}{n_f} - \frac{1}{n_i} \right)} \quad \text{For } n_f > n_i$$

Note: The energy, frequency, and wavelength of emitted or absorbed light should always be a **positive** number! The words absorption and emission indicate whether energy is being lost or gained.

Summary for absorption/emission

The Rydberg formula can be used to calculate the frequency (and also the E or λ , using $E = h\nu$ or $\lambda = c/\nu$) of light emitted or absorbed by any one-electron atom or ion. Note: adding Z makes the equation generic for any one-electron atom or ion. For hydrogen, Z=1.

$$\nu = \frac{Z^2 R_H}{h} \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

for $n_i > n_f$

Electron _____ energy

$$\nu = \frac{Z^2 R_H}{h} \left(\frac{1}{n_i^2} - \frac{1}{n_f^2} \right)$$

for $n_f > n_i$

Electron _____ energy

Overall: The Schrödinger equation correctly predicts the hydrogen-atom emission spectrum, showing that the binding energies derived from the Schrödinger equation are in agreement with experiment. The Schrödinger equation also tells us about wavefunctions (orbitals) – that topic is next.

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