# MASSACHUSETTS INSTITUTE OF TECHNOLOGY 

Physics Department

## Solutions to Problem Set \#7

Problem 1: Free Expansion of a Gas
a) No work is done, so $\Delta W=0$. No heat enters the gas so $\Delta Q=0$.

Thus $\Delta E=\Delta W+\Delta Q=0$. The internal energy is conserved.
b) $E(T, V)$ is a state function; compare it before and after expansion in equilibrium situations.

$$
\begin{aligned}
d E & =T d S-P d V=\underbrace{T\left(\frac{\partial S}{\partial T}\right)_{V}}_{C_{V}} d T+\left(T\left(\frac{\partial S}{\partial V}\right)_{T}-P\right) d V \\
C_{V} & =\frac{3}{2} N k \\
\left(\frac{\partial S}{\partial V}\right)_{T} & =\left(\frac{\partial P}{\partial T}\right)_{V}=\frac{N k}{V-b N} \quad \text { using a Maxwell relation } \\
d E & =\frac{3}{2} N k d T+a\left(\frac{N}{V}\right)^{2} d V \\
E & =\frac{3}{2} N k T-\frac{a N^{2}}{V}+\text { constant }
\end{aligned}
$$

When equating $E$ before and after the constant cancels out.

$$
\begin{aligned}
\frac{3}{2} N k T_{i}-\frac{a N^{2}}{\left(V_{0} / 3\right)} & =\frac{3}{2} N k T_{f}-\frac{a N^{2}}{V_{0}} \\
\frac{3}{2} N k\left(T_{f}-T_{i}\right) & =-\frac{2 a N^{2}}{V_{0}} \\
T_{f} & =T_{i}-\frac{4}{3} \frac{a}{k}\left(\frac{N}{V_{0}}\right)
\end{aligned}
$$

The gas cools!

Problem 2: Use of a Carnot Cycle
a) Carnot Cycle $\Rightarrow d S_{H}=-d S_{C} \Rightarrow d Q_{H} / T_{H}=-\phi Q_{C} / T_{C}$. Use $d Q=C_{0} d T$.

$$
\begin{aligned}
\Delta S_{b o d y} & =-\Delta S_{b o d y} 2 \\
\int_{T_{H}}^{T_{F}} \frac{C_{0}}{T} d T & =-\int_{T_{C}}^{T_{F}} \frac{C_{0}}{T} d T \\
\ln \frac{T_{F}}{T_{H}} & =-\ln \frac{T_{F}}{T_{C}}=\ln \frac{T_{C}}{T_{F}} \\
\frac{T_{F}}{T_{H}} & =\frac{T_{C}}{T_{F}} \\
T_{F} & =\underline{\sqrt{T_{H} T_{C}}}
\end{aligned}
$$

b)

$$
\begin{aligned}
\Delta W_{\text {out }} & =\left|\Delta Q_{H}\right|-\left|\Delta Q_{C}\right| \\
W_{\text {out }} & =\int_{T_{F}}^{T_{H}} C_{0} d T-\int_{T_{C}}^{T_{F}} C_{0} d T=C_{0}\left[\left(T_{H}-T_{F}\right)-\left(T_{F}-T_{C}\right)\right] \\
& =C_{0}\left(T_{H}-2 T_{F}+T_{C}\right)=C_{0}\left(T_{H}-2 \sqrt{T_{H} T_{C}}+T_{C}\right) \\
& =\underline{C_{0}\left(\sqrt{T_{H}}-\sqrt{T_{C}}\right)^{2}>0}
\end{aligned}
$$

## Problem 3: Cooling Liquid Helium

a) Since the system is thermally isolated and no work is done in the process, the heat gained by the salt must equal the heat lost by the liquid.

$$
\begin{aligned}
\Delta Q_{S} & =-\Delta Q_{L} \\
\int_{T_{0}}^{1 / 2} b T^{-2} d T & =-\int_{1}^{1 / 2} a T^{3} d T \\
-b \int_{T_{0}}^{1 / 2} T^{-1} & =-\frac{a}{4} \int_{1}^{1 / 2} T^{4} \\
\left(2-\frac{1}{T_{0}}\right) & =\frac{1}{4} \frac{a}{b}\left(\frac{1}{16}-1\right)=\frac{1}{4} \frac{a}{b} \frac{15}{16}=\frac{1}{4} \frac{128}{15} \frac{15}{16}=-2 \\
1 / T_{0} & =4 \rightarrow \underline{T_{0}=1 / 4}
\end{aligned}
$$

b) Entropy is a state function, so one may compute its change by assuming the process was quasi-static. We will use $d S=\not \subset Q / T=C d T / T$.

$$
\begin{aligned}
\Delta S & =\Delta S_{S}+\Delta S_{L} \\
& =\int_{T_{0}}^{1 / 2} b T^{-3} d T+\int_{1}^{1 / 2} a T^{2} d T \\
& =-\frac{b}{2} \int_{1 / 4}^{1 / 2} T^{-2}+\frac{a}{3} \int_{1}^{1 / 2} T^{3} \\
& =-\frac{b}{2} \underbrace{(4-16)}_{-12}+\frac{a}{3} \underbrace{\left(\frac{1}{8}-1\right)}_{-7 / 8}=6 b-(7 / 24) a
\end{aligned}
$$

## Problem 4: Torsional Pendulum

a) First find the Hamiltonian.

$$
\mathcal{H}=T+V+\frac{1}{2} I \dot{\theta}^{2}+\frac{1}{2} K\left(\theta-\theta_{0}\right)^{2}
$$

Then use the canonical ensemble expression for the probability density.

$$
\begin{aligned}
p(\theta, \dot{\theta}) & \propto \exp \left[-\frac{\mathcal{H}}{k T}\right] \\
& \propto \exp \left[-\frac{\dot{\theta}^{2}}{2(k T / I)}\right] \exp \left[-\frac{\left(\theta-\theta_{0}\right)^{2}}{2(k T / K)}\right]
\end{aligned}
$$

Notice two important features of this result. First, the probability density factors, $p(\theta, \dot{\theta})=$ $p(\theta) p(\dot{\theta})$, so $\theta$ and $\dot{\theta}$ are statistically independent. Second, the dependence on both $\theta$ and $\dot{\theta}$ has the Gaussian form. In particular, $p(\theta)$ is Gaussian with mean $\theta_{0}$ and variance $\sigma_{\theta}^{2}=k T / K$. Therefore,

$$
<\left(\theta-\theta_{0}\right)^{2}>^{1 / 2}=\sqrt{\frac{k T}{K}} .
$$

b) Since $\theta$ and $\dot{\theta}$ are statistically independent, $\langle\theta \dot{\theta}>=<\theta><\dot{\theta}>$. By inspection, $p(\dot{\theta})$ is a zero-mean Gaussian, so $\langle\dot{\theta}\rangle=0$ which leads to $\leq \dot{\theta}\rangle=0$.

Problem 5: The Hydrogen Atom
a)

$$
\mathcal{H}\left|n, l, m>=-\frac{A}{n^{2}}\right| n, l, m>
$$

The lowest energy, $-A$, corresponds to $\mid 1,0,0>$ and is non-degenerate. The next lowest energy, $-A / 4$, is four fold degenerate:

$$
|2,0,0>, \quad| 2,1,1>, \quad \mid 2,1,0>, \quad \text { and } \mid 2,1,-1>
$$

The ratio of the number of atoms in the first excited energy level to the number in the ground state depends on both the energies and the degeneracies.

$$
\frac{N(-A / 4)}{N(-A)}=\frac{4 \exp [A / 4 k T]}{\exp [A / k T]}=4 \exp [-(3 / 4) A / k T]
$$

Using the conversion factor $1 \mathrm{meV}=11.6 \mathrm{~K}$ we find that $13.6 \mathrm{eV}=1.58 \times 10^{5} \mathrm{~K}$. Evaluating the above ratio gives $4.8 \times 10^{-170}$ at 300 K and $1.6 \times 10^{-51}$ at 1000 K .
b) The degeneracy of the $n^{\text {th }}$ energy level is

$$
1+3+5+\cdots+(2 n-1)=n^{2}
$$

The partition function for a single atom, neglecting the unbound states, is

$$
Z=\sum_{\text {states } i} \exp \left[-\epsilon_{i} / k T\right]=\sum_{n=1}^{\infty} n^{2} \exp \left[\alpha / n^{2}\right]
$$

where $\alpha \equiv A / k T$. Since $\alpha>0$, it follows that $\exp \left[\alpha / n^{2}\right]>1$ for all $n$. Using this we can set a lower bound for $Z$, but $Z$ diverges since the lower bound diverges.

$$
Z>\sum_{n=1}^{\infty} n^{2} \quad \text { which diverges }
$$

c) The Coulomb potential is a mathematical oddity in that it produces an infinite number of bound states with energies less than zero. This situation if modified in the real world by the presence of walls (consider the energy levels of a particle in a box) or by the presence of other atoms. The existence of hydrogen atoms in the interstellar medium, on the other hand, probably has more to do with the absence of excitation mechanisms (non-equilibrium) than with the presence of neighboring atoms.

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