MASSACHUSETTS INSTITUTE OF TECHNOLOGY

Physics Department

8.044 Statistical Physics I

Spring Term 2013

Solutions, Exam #2

Problem 1 (30 points) Entropy of a Surface Film

 γ and C_A are given in terms of T and A so it is reasonable to choose T and A as the variables in which to expand the entropy.

$$dS = \left(\frac{\partial S}{\partial T}\right)_{A} dT + \left(\frac{\partial S}{\partial A}\right)_{T} dA$$
$$C_{A} \equiv \left.\frac{dQ}{dt}\right|_{A} = T \left(\frac{\partial S}{\partial T}\right)_{A} \Rightarrow \left(\frac{\partial S}{\partial T}\right)_{A} = \frac{C_{A}}{T} = \frac{Nk_{B}}{T} + \frac{Nk_{B}T}{T_{0}^{2}}$$

To find $(\partial S/\partial A)_T$ use a Maxwell Relation. You may either use the magic square or derive the required relation as follows.

$$F \equiv U - TS$$
$$dF = -SdT + \gamma dA$$

cross derivatives of the prefactors of the differentials are equal

$$\left(\frac{\partial S}{\partial A}\right)_T = -\left(\frac{\partial \gamma}{\partial T}\right)_A = \frac{Nk_B}{A - bN}$$

Substituting in these results gives

$$dS = \left(\frac{Nk_B}{T} + \frac{Nk_BT}{T_0^2}\right) dT + \left(\frac{Nk_B}{A - bN}\right) dA$$

$$S = Nk_B \ln T + \frac{1}{2}Nk_B \left(\frac{T}{T_0}\right)^2 + f(A)$$

$$\left(\frac{\partial S}{\partial A}\right)_T = f'(A) = \frac{Nk_B}{A - bN} \Rightarrow f(A) = Nk_B \ln(A - bN) + c$$

$$S(T, A) = Nk_B \ln T + \frac{1}{2}Nk_B \left(\frac{T}{T_0}\right)^2 + Nk_B \ln(A - bN) + c$$

[Note: One can make the arguments of the logs dimensionless by distributing part of the additive constant c among the various other terms.]

$$S(T,A) = Nk_B \ln(T/T_1) + \frac{1}{2}Nk_B \left(\frac{T}{T_0}\right)^2 + Nk_B \ln((A-bN)/A_1) + c'$$

Problem 2 (40 points) Crystal Field Splitting

a)

$$Z_1 = 1 + 2 \exp[-\Delta/k_B T]$$

$$<\epsilon>$$
 = $\sum_{\text{states}} \epsilon_{\text{state}} p(\text{state}) = \frac{2\Delta \exp[-\Delta/k_B T]}{1 + 2\exp[-\Delta/k_B T]} = 2\Delta \frac{1}{\exp[\Delta/k_B T] + 2}$
 $U(T, N) = N < \epsilon >= 2\Delta N \frac{1}{\exp[\Delta/k_B T] + 2}$

b) At T = 0 only the non-degenerate ground state is occupied. $\underline{S(T = 0, N) = k_B N \ln(1) = 0}$. As $T \to \infty$, all three states are equally probable. $\underline{S(T, N) \to k_B N \ln(3)}$.

c)

$$C_{V} = \left(\frac{\partial U}{\partial T}\right)_{V} = 2\Delta N \frac{d}{dT} \left(\frac{1}{\exp[\Delta/k_{B}T]+2}\right) \qquad C_{V} \qquad \text{energy gap behavior}$$

$$= 2\Delta N \frac{(\Delta/k_{B}T^{2}) \exp[\Delta/k_{B}T]}{(\exp[\Delta/k_{B}T]+2)^{2}} \qquad \qquad I/T^{2}$$

$$= 2Nk_{B} \left(\frac{\Delta}{k_{B}T}\right)^{2} \frac{\exp[\Delta/k_{B}T]}{(\exp[\Delta/k_{B}T]+2)^{2}} \qquad \qquad T$$

d)

$$F(T,N) = -k_B T \ln Z = -Nk_B T \ln Z_1 = -Nk_B T \ln(1 + 2\exp[-\Delta/k_B T])$$

$$P(T,N) = -\left(\frac{\partial F}{\partial V}\right)_{T} = -\left(\frac{\partial F}{\partial \Delta}\right)_{T} \frac{d\Delta}{dV} = \gamma \left(\frac{\Delta}{V}\right) \left(\frac{\partial F}{\partial \Delta}\right)_{T}$$
$$= -Nk_{B}T\gamma \left(\frac{\Delta}{V}\right) \frac{1}{Z_{1}} \left(\frac{-2}{k_{B}T}\right) \exp[-\Delta/k_{B}T]$$
$$= 2N\gamma \left(\frac{\Delta}{V}\right) \frac{\exp[-\Delta/k_{B}T]}{1+2\exp[-\Delta/k_{B}T]}$$
$$= \frac{\left(\frac{\gamma}{V}\right) 2N\Delta \frac{1}{\exp[\Delta/k_{B}T]+2}}{\exp[\Delta/k_{B}T]+2} = \gamma \frac{U}{V}$$

Problem 3 (30 points) Heating a Shell

a) For the shell,

$$P_{\text{in}} = 4\pi r^2 \sigma T_H^4$$

$$P_{\text{out}} = 4\pi R^2 \sigma T_S^4$$

$$P_{\text{out}} = P_{\text{in}} \Rightarrow r^2 T_H^4 = R^2 T_S^4 \rightarrow \underline{T_S = T_H \sqrt{\frac{r}{R}}}$$

b)

$$e(\omega, T)_{\text{heater}} = (1) \left(\frac{1}{4}\right) c u(\omega, T_H)$$

$$P_{\text{in}} = (4\pi r^2) \left(\frac{c}{4}\right) \int_0^{\omega_0} \left(\frac{k_B T_H}{\pi^2 c^3}\right) \omega^2 d\omega$$

$$= \frac{r^2 k_B T_H}{\pi c^2} \int_0^{\omega_0} \omega^2 d\omega = \frac{k_B \omega_0^3}{3\pi c^2} r^2 T_H$$

Note that the power is coming from the central object (not from the shell) and from its surface (not volume). Thus this result is proportional to r^2 .

c)

$$e(\omega, T)_{\text{shell}} = \alpha(\omega) \left(\frac{1}{4}\right) c u(\omega, T_S)$$

$$P_{\text{out}} = (4\pi R^2) \left(\frac{c}{4}\right) \int_0^{\omega_0} \left(\frac{k_B T_S}{\pi^2 c^3}\right) \omega^2 d\omega = \frac{k_B \omega_0^3}{3\pi c^2} R^2 T_S$$

$$P_{\text{out}} = P_{\text{in}} \Rightarrow \underline{T_S = T_H \left(\frac{r}{R}\right)^2}$$

This is an example of a poor absorber being a poor emitter (Kirchoff's law, on the information sheet). The shell does not absorb beyond ω_0 , thus it does not radiate beyond ω_0 .

MIT OpenCourseWare http://ocw.mit.edu

8.044 Statistical Physics I Spring 2013

For information about citing these materials or our Terms of Use, visit: http://ocw.mit.edu/terms.