3. Hint for Problem 3, Carroll problem 7.1.

Carroll 7.1 asks us to vary the Lagrangian

$$
\mathcal{L}=\frac{1}{2}\left[\left(\partial_{\mu} h^{\mu \nu}\right)\left(\partial_{\nu} h\right)-\left(\partial_{\mu} h^{\rho \sigma}\right)\left(\partial_{\rho} h_{\sigma}^{\mu}\right)+\frac{1}{2} \eta^{\mu \nu}\left(\partial_{\mu} h^{\rho \sigma}\right)\left(\partial_{\nu} h_{\rho \sigma}\right)-\frac{1}{2} \eta^{\mu \nu}\left(\partial_{\mu} h\right)\left(\partial_{\nu} h\right)\right]
$$

to construct the Einstein tensor, Eq. (7.8) of Carroll.
If you vary this Lagrangian in the most straightforward way possible, you will probably find that you get almost the correct Einstein tensor - you should get Eq. (7.8), but with the first two terms replaced with 2 times the first term. In other words, you don't get the symmetrization on $\mu$ and $\nu$ that should be obtained.
What is going on here? The issue is that the Lagrangian doesn't "know", a priori, that the tensor $h_{\mu \nu}$ is symmetric. This has a strong impact on the second term of the Lagrangian - it should be symmetric with respect to exchange of the indices $\rho$ and $\sigma$, but isn't unless you somehow build in the knowledge we have of this symmetry.
There are two simple ways to address this:
a. Rewrite the Lagrangian to force this symmetrization:

$$
\left(\partial_{\mu} h^{\rho \sigma}\right)\left(\partial_{\rho} h^{\mu}{ }_{\sigma}\right) \rightarrow \frac{1}{2}\left[\left(\partial_{\mu} h^{\rho \sigma}\right)\left(\partial_{\rho} h^{\mu}{ }_{\sigma}\right)+\left(\partial_{\mu} h^{\rho \sigma}\right)\left(\partial_{\sigma} h^{\mu}{ }_{\rho}\right)\right],
$$

or
b. Make sure that, in our variation, this symmetry is enforced. The way I did this was to note that I should have

$$
\begin{aligned}
\frac{\delta\left(\partial_{\mu} h_{\rho \sigma}\right)}{\delta\left(\partial_{\gamma} h_{\alpha \beta}\right)} & =\frac{\delta\left(\partial_{\mu} h_{\rho \sigma}\right)}{\delta\left(\partial_{\gamma} h_{\beta \alpha}\right)} \\
& =\frac{1}{2}\left[\frac{\delta\left(\partial_{\mu} h_{\rho \sigma}\right)}{\delta\left(\partial_{\gamma} h_{\alpha \beta}\right)}+\frac{\delta\left(\partial_{\mu} h_{\rho \sigma}\right)}{\delta\left(\partial_{\gamma} h_{\beta \alpha}\right)}\right] \\
& =\frac{1}{2}\left[\delta^{\gamma}{ }_{\mu} \delta^{\alpha}{ }_{\rho} \delta^{\beta}{ }_{\sigma}+\delta^{\gamma}{ }_{\mu} \delta^{\alpha}{ }_{\sigma} \delta^{\beta}{ }_{\rho}\right]
\end{aligned}
$$

It shouldn't be too difficult to convince yourself that these metbds are in fact equivalent.

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### 8.962 General Relativity

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